

Repercussions of a geometric theory of flavor and the 126 GeV SMS

Amarjit Soni, HET, BNL

Scalars 2013, Warsaw 9/13/13

For further details....

- Talk is primarily based on arXiv: 1303.5056
- see also Davoudiasl, McElmurry and A.S. arXiv:1206.4062
- See also talks at
- FPCP 2012, Hefei China, May 2012
- ICHEP 2012 Melbourne (July 2012)
- Solvay workshop on “Facing the Scalar Sector” , Brussels, May 29-31, 2013

Main point

- Flavor constraints for long have been telling us that scale of new physics is not near 1 TeV.
- Recent discovery of scalar (126 GeV) with SM-like properties, *completely independently*, tells us that the scale of new physics is not likely ~ 1 TeV
- So what is it?
- What are its consequences?

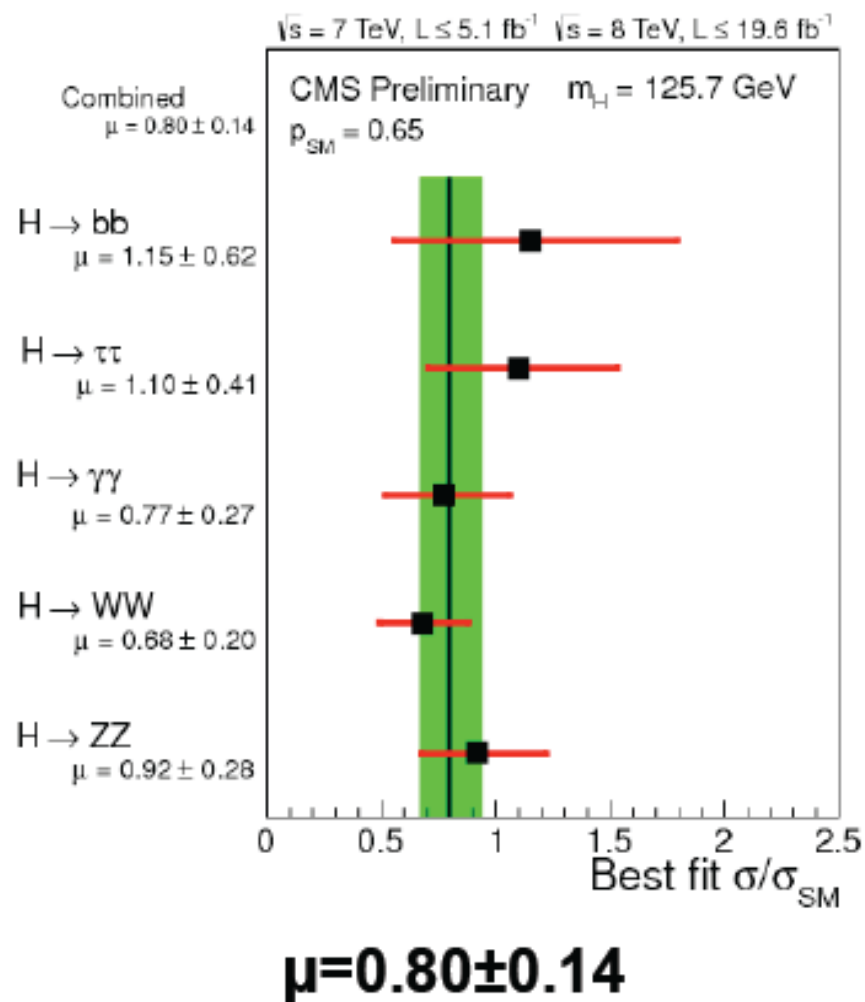
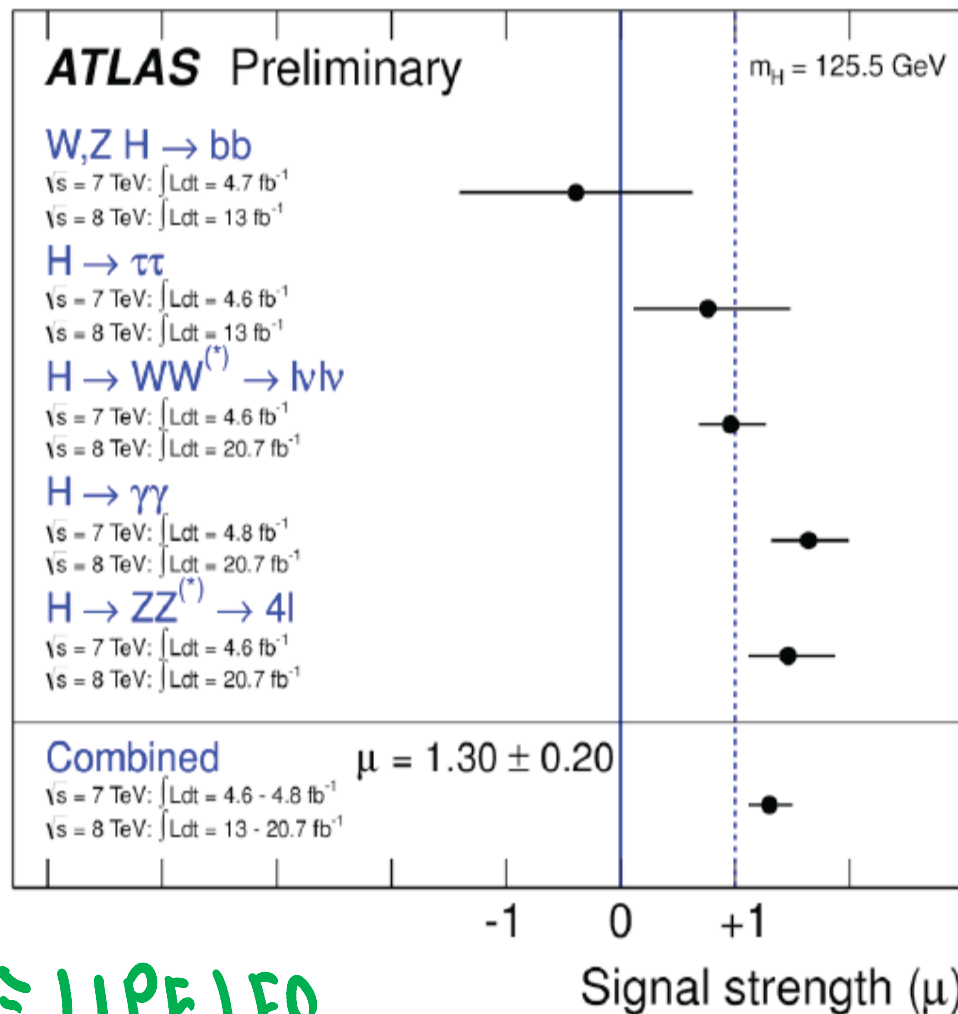
[From the perspective of a warp theory of flavor]

Outline

- Works but
- Scale of NP=> future directions
- From the vantage point of a geometric theory of flavor scale was not expected to be less than around 10 TeV*
- summary

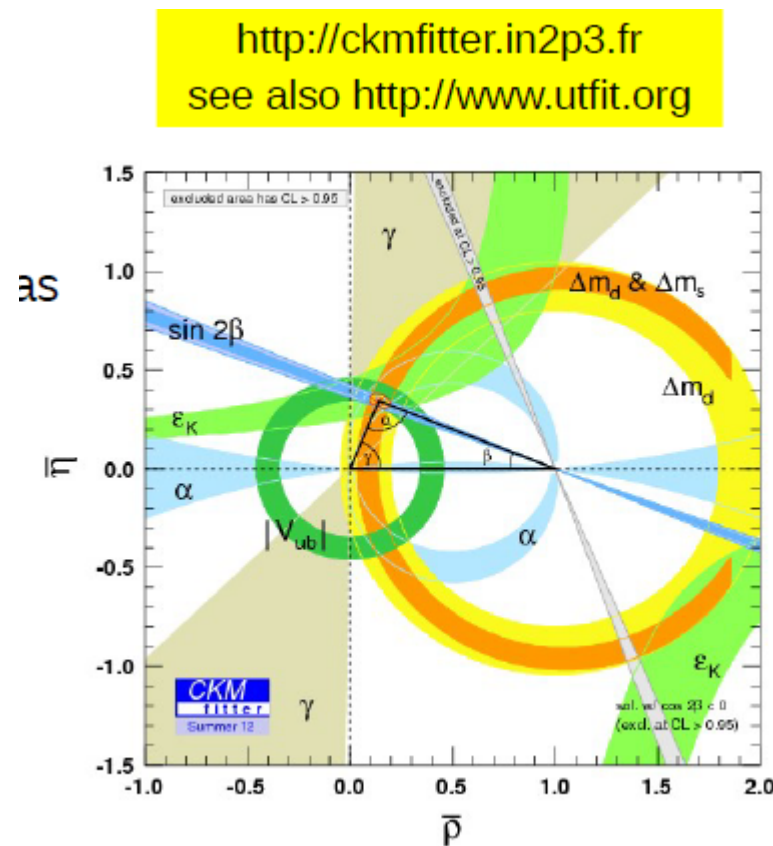
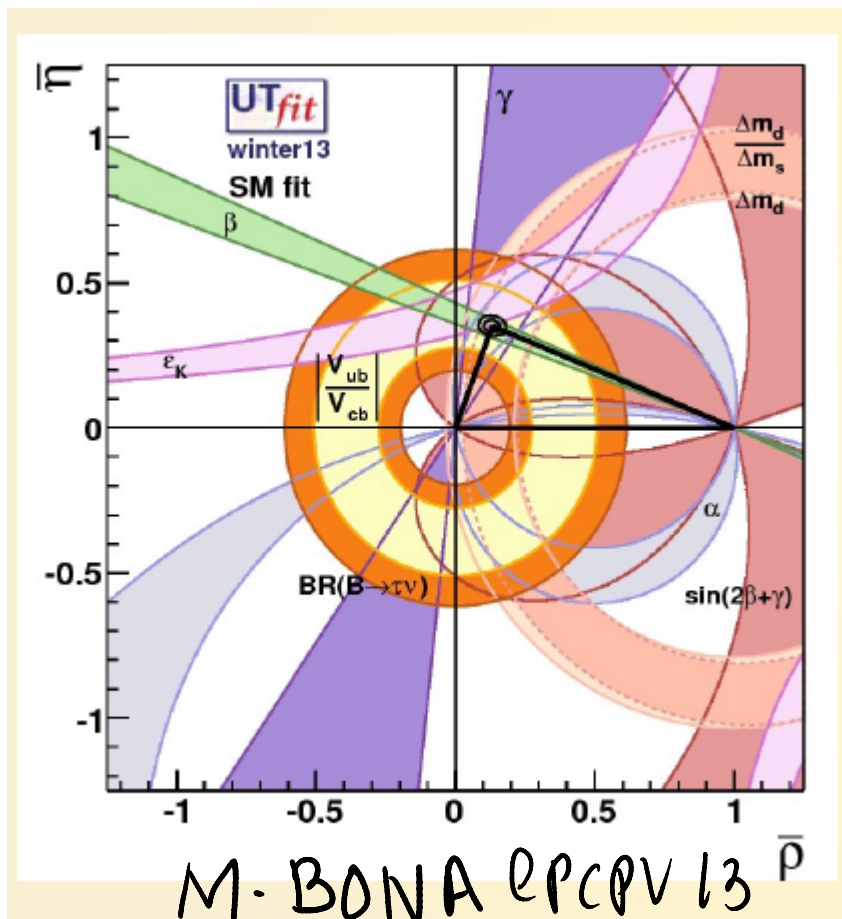
* Flavor is not a footprint and, in particular flavor-alignment is a serious issue

FITS LIKE A GLOVE!
[OR DOES IT?]



E. LIPELES
PBF2013

S. BOSE PBF2013



T. GERSHON
CPCF2013

SM-CKM paradigm works rather well.

No glaring discrepancy

OTDH tests only $\sim 10-15\%$ accuracy

Is Nature Unnatural?

Decades of confounding experiments have physicists considering a startling possibility: The universe might not make sense.

by: [Natalie Wolchover](#)

May 24, 2013

[email](#) [print](#)



Is the universe natural or do we live in an atypical bubble in a multiverse? Recent results at the Large Hadron Collider have forced many physicists to confront the latter possibility. (Illustration: Giovanni Villadoro)

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A lesson from history (I)

"A special search at Dubna was carried out by E. Okonov and his group. They did not find a single $K_L \rightarrow \pi^+ \pi^-$ event among 600 decays into charged particles [12] (Anikira et al., JETP 1962). At that stage the search was terminated by the administration of the Lab. The group was unlucky."

-Lev Okun, "The Vacuum as Seen from Moscow"

1964: $BF = 2 \times 10^{-3}$

A failure of imagination ? Lack of patience ?

CHRISTENSEN,
CANNON, FITCH
& TURLAY
BNL 1964

DRAWING STRONG CONCLUSIONS
BASED ON 20% Tests is
TOO RISKY!

**SHOULD WE BE SHOCKED TO FIND THAT
THE SCALE OF NEW PHYSICS IS NOT ~ 1
TEV & APPEARS TO BE HIGHER?**

Outstanding Th.puzzles of our times

- Hierarchy puzzle**

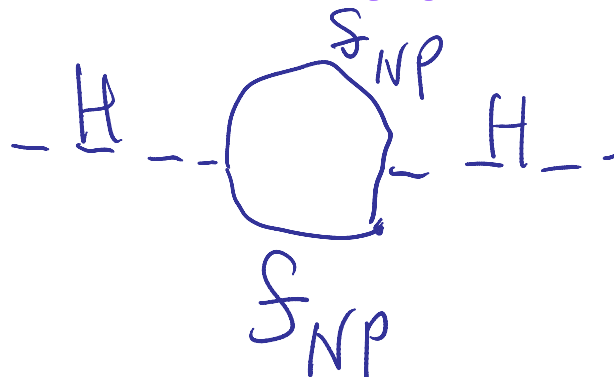


Diagram: A loop diagram with two external Higgs lines (H) and two internal scalar lines labeled S_{NP} .

$$-H--\text{loop}(S_{NP})--H- \sim \frac{g_{NP}^2}{16\pi^2} \Lambda_{NP}^2 \Rightarrow \Lambda_{NP} \lesssim \text{TeV}$$

to avoid fine tuning m_H

- Flavor puzzle**

$\Delta f_{\text{flavor}} = 2$ e.g.

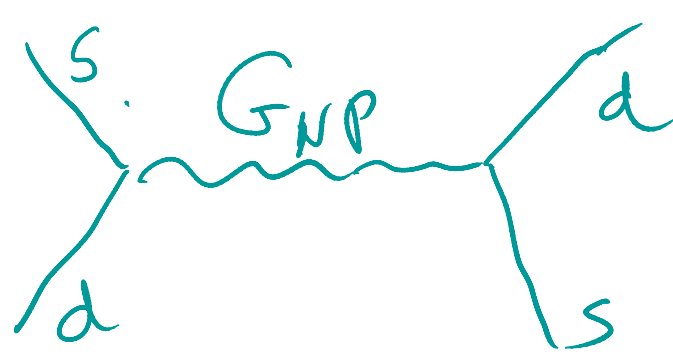
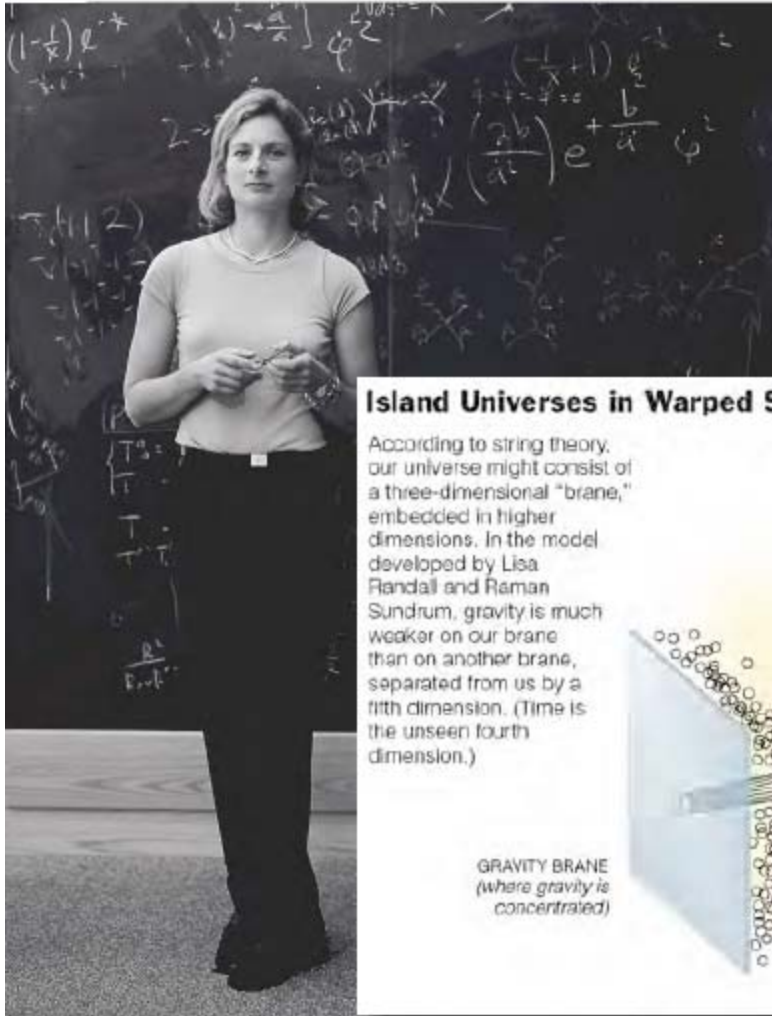



Diagram: A tree-level exchange diagram with two quark lines (s and d) and a scalar NP line labeled G_{NP} .

$$\sim \frac{g_{NP}^2}{\Lambda_{NP}^2} \Rightarrow \Lambda_{NP} \gtrsim 10^3 \text{ TeV}$$

to avoid constraint from Δm_K

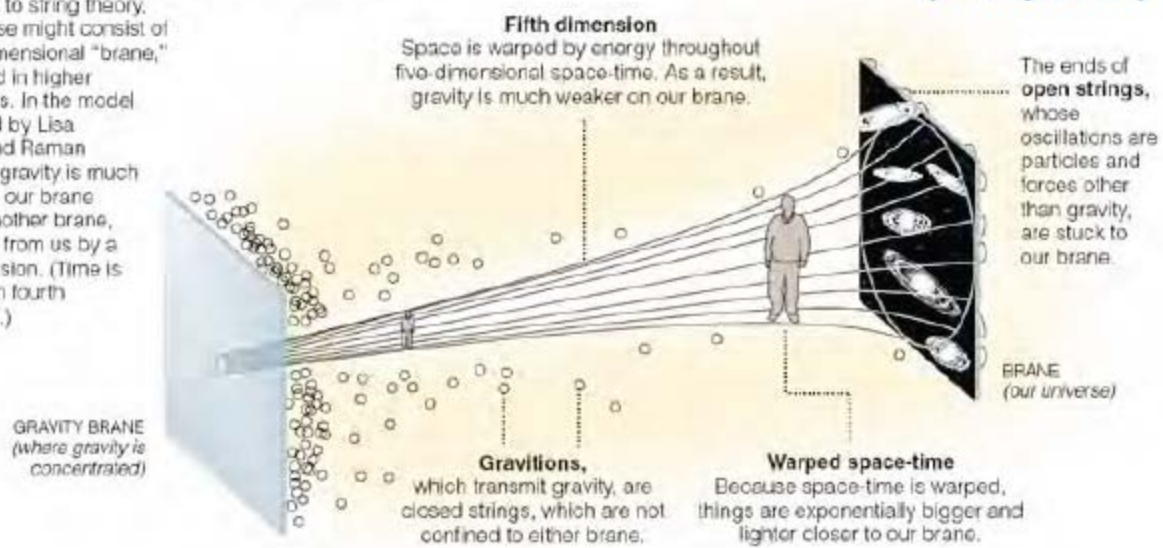
The Randall-Sundrum (RS) idea

Island Universes in Warped Space-Time

According to string theory, our universe might consist of a three-dimensional "brane," embedded in higher dimensions. In the model developed by Lisa Randall and Raman Sundrum, gravity is much weaker on our brane than on another brane, separated from us by a fifth dimension. (Time is the unseen fourth dimension.)

(Wikipedia)



GRAVITY BRANE
(where gravity is concentrated)

Fifth dimension
Space is warped by energy throughout five-dimensional space-time. As a result, gravity is much weaker on our brane.

Gravitons,
which transmit gravity, are closed strings, which are not confined to either brane.

Warped space-time
Because space-time is warped, things are exponentially bigger and lighter closer to our brane.

BRANE
(our universe)

The ends of **open strings**, whose oscillations are particles and forces other than gravity, are stuck to our brane.

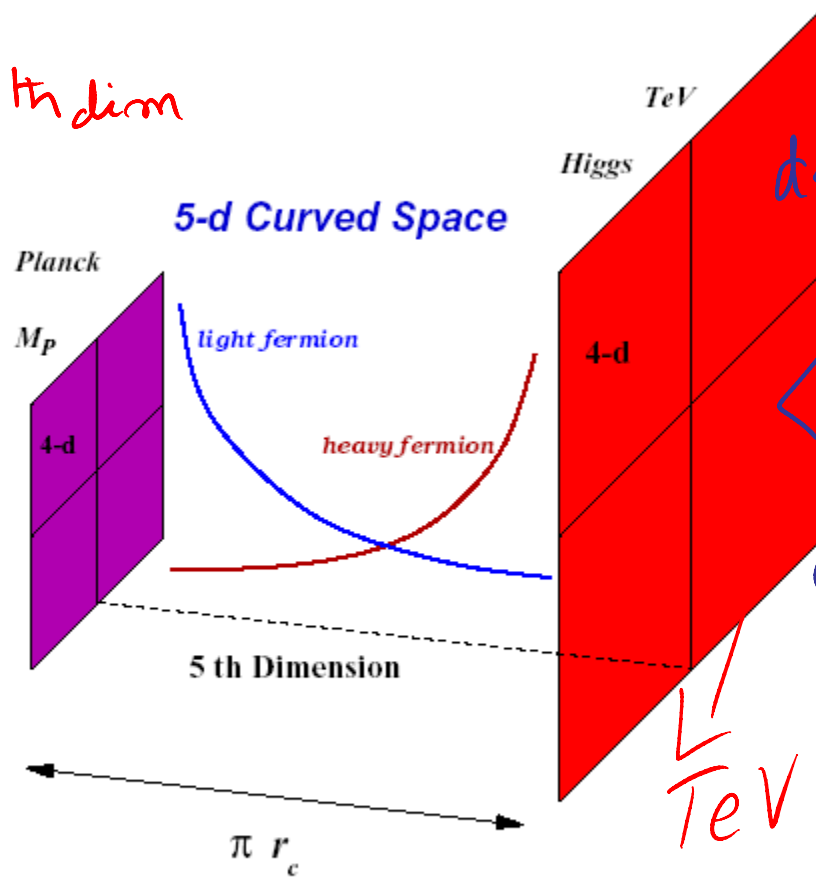
[stolen from Newt]

INSIGHTS FROM A GEOMETRIC THEORY OF FLAVOR

RANDALL+SUNDROM '99

[FIG BY
H DAVOUDI@SL]

Points along 5th dim
correspond to
diff. eff.
4d scale!



$$ds^2 = e^{-2\sigma} \eta_{\mu\nu} dx^\mu dx^\nu - r_c^2 d\psi^2$$

$$\langle H_4 \rangle = e^{-6} \langle H_5 \rangle$$

$$G = \frac{1}{2} r_c \pi$$

$$\sim \frac{1}{12}$$

$$M_P$$

Figure 1: Warped geometry with flavor from fermion localization. The Higgs field resides on the TeV-brane. The size of the extra dimension is $\pi r_c \sim M_P^{-1}$.

Simultaneous resolution to hierarchy and flavor puzzles

Fermion “geography” (localization) naturally explains:

Grossman&Neubert; Gherghetta&Pomarol; Davoudiasl, Hewett & Rizzo

- Why they are light (or heavy)
 - FCNC for light quarks are severely suppressed automatically
 - RS-GIM MECHANISM (Agashe, Perez, AS'04) **flavor changing transitions though at the *tree level* (resulting from rotation from interaction to mass basis) are suppressed roughly to the same level as the loop in SM \Rightarrow CKM mixings (& mass) hierarchy.**
 - **O(1) CP ubiquitous;.....nedm, in fact ALL DIR-CP [ε'/ε , γ , $\Delta ACP(B \Rightarrow K\pi)$, $\Delta(\sin 2\beta)$; $S[B \Rightarrow K^* \rho\gamma]$; $\Delta ACP(D)$..] are an exceedingly important path to BSM-phase and new physics**
 - Most flavor violations are driven by the top
- > ENHANCED $t \rightarrow cZ(h)$ A VERY IMPORTANT “GENERIC” PREDICTION..Agashe, Perez, AS'06

$$E_K, \Delta m_K : 10^3 \text{ TeV} \Rightarrow RS_{Fl} \sim 10 \text{ TeV}$$

EXTENSIVE STUDIES by BURAS et al and by NEUBERT et al

Localization parameters of the 3-families of quarks

$$\begin{array}{lll} c_{Q_1} = -0.579, & c_{Q_2} = -0.517, & c_{Q_3} = -0.473 \\ c_{u_1} = -0.742, & c_{u_2} = -0.558, & c_{u_3} = +0.339 \\ c_{d_1} = -0.711, & c_{d_2} = -0.666, & c_{d_3} = -0.553 \end{array}$$

Table from
M. Neubert
@Moriond09

⇒ masses of the 6 quarks in RS!

The bulk profile of a fermionic zero mode depends strongly on its bulk mass parameter c_Ψ . In case of a left-handed zero mode $\Psi_L^{(0)}$ it is given by [2, 4]

$$f_L^{(0)}(y, c_\Psi) = \sqrt{\frac{(1 - 2c_\Psi)kL}{e^{(1-2c_\Psi)kL} - 1}} e^{-c_\Psi ky} \quad (2.6)$$

with respect to the warped metric. Therefore, for $c_\Psi > 1/2$ the fermion $\Psi_L^{(0)}$ is localised towards the UV brane and exponentially suppressed on the IR brane, while for $c_\Psi < 1/2$

M. BLANK E et al JHEP'09

REMINISCENT OF CHENG-SHER ANZATS

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$$f_{ij} = \lambda_{ij} \sqrt{m_i m_j} / v$$

Minimal tuning ansatz & naturalness

- RS_{Fl} is the scale at which all flavor constraints are satisfied....
- This scale seems to be set by Kaon mixings and is roughly around 10 TeV. This is high enough that EWPC are automatically satisfied.
- While scale of new physics could be bigger than this, its likely that it is close to RS_{Fl} so as to minimize the degree of tuning needed....

$$v^2 / [RS_{Fl}]^2 \sim 10^{-3} \gg \gg [v/M_{pl}]^2$$

Natuaralness is not at stake at least not yet!

B-Factory Signals for a Warped Extra DimensionKaustubh Agashe,^{1,*} Gilad Perez,^{2,†} and Amarjit Soni^{3,‡}¹*Department of Physics and Astronomy, Johns Hopkins University, Baltimore, Maryland 21218-2686, USA*²*Theoretical Physics Group, Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA*³*Brookhaven National Laboratory, Upton, New York 11973, USA*

(Received 23 June 2004; published 10 November 2004)

m_{KK} ~ 3 TeV

We study predictions for B physics in a class of warped extra dimension models recently introduced, where few (~ 3) TeV Kaluza-Klein masses are consistent with electroweak data due to custodial symmetry. As in the standard model (SM), flavor violations arise due to the heavy top quark leading to striking signals: (i) New physics contributions to $\Delta F = 2$ transitions are comparable to the SM, so the success of the SM unitarity triangle fit is a “coincidence.” Thus, clean extractions of unitarity angles are likely to be affected, in addition to $O(1)$ deviation from the SM prediction in B_s mixing. (ii) $O(1)$ deviation from various SM predictions for $B \rightarrow X_s l^+ l^-$. (iii) Large mixing-induced CP asymmetry in radiative B decays. Also, the neutron electric dipole moment is roughly 20 times larger than the current bound so that this framework has a “ CP problem.”

TABLE I. Contrasting signals from RS1 with the SM.

	Δm_{B_s}	$S_{B_s \rightarrow \psi\phi}$	$S_{B_d \rightarrow \phi K_s}$	$Br[b \rightarrow sl^+l^-]$	$S_{B_{d,s} \rightarrow K^*, \phi\gamma}$	$S_{B_{d,s} \rightarrow \rho, K^*\gamma}$
RS1	$\Delta m_{B_s}^{\text{SM}}[1 + O(1)]$	$O(1)$	$\sin 2\beta \pm O(0.2)$	$Br^{\text{SM}}[1 + O(1)]$	$O(1)$	$O(1)$
SM	$\Delta m_{B_s}^{\text{SM}}$	λ_c^2	$\sin 2\beta$	Br^{SM}	$\frac{m_s}{m_b}(\sin 2\beta, \lambda_c^2)$	$\frac{m_d}{m_b}(\lambda_c^2, \sin 2\beta)$

In our '04 papers 3 TeV scale was an unfortunate oversight

- FC KK-glu exchanges give rise to LR currents
- LR currents cause enhanced Kaon mixings.....
- [Beall, Bander, AS, PRL '82]....
- C also M.Bona et al [UTFit] arXiv:0707.0636
- Agashe, Papucci, Perez and Pirjol, hep-ph/0509117
- Gedalia, Isidori and Perez, arXiv: 0905.3264

$$\Rightarrow m_{KK} \sim 10 \text{ TeV}$$

EWPC/Little Hierarchy

- Imposing custodial symmetry, by enlarging $SU2 \Rightarrow SU2 \times SU2$ and by adding additional fermions, $m_{KK} > \sim 3 \text{ TeV}$ [see Agashe, Delgado, May & Sundrum'03]
- Once $RS_{FI} > 10 \text{ TeV}$ is imposed, EWPC likely automatically satisfied; there is now **“Little Hierarchy”** i.e. **tuning $O(10^{-3})$** is somewhat worse but the setup is more economical & simpler; perhaps the experimental **indications to date are that this simplicity is preferred.**
- Admittedly, this is not the last word yet; experiment as always is the decider and LHC(~ 14) will provide an important clue

**So far we got from flavor
perspective**

**Next we look at it from the
perspective of SMS (126 GeV)**

Higgs is SM-like =>

- Light SM-like Higgs strengthens case for $m_{KK} > \sim 10$ TeV in warped framework

M. Carena et al, 1204.0008

Davoudiasl, McElmurry, A. S. 1206.4062

1. With $m_{KK} \sim 10$ TeV resulting set up is simpler and economical as then may NOT need to enforce custodial symmetry
(which requires introducing more dofs....see Agashe et al...)
2. With $m_{KK} \sim 10$ TeV, tuning is somewhat worse but has the advantage of a more economical theoretical framework and may well be preferable.
3. At LHC at best only radion may be accessible.

Important observables & some expectations

- **For The Intensity Frontier**
- $\mu \rightarrow e \gamma$ within factors of $O(\text{few})$ close to Expt bound $< 6 \times 10^{-26}$ e-cm
- Null tests extremely important ; Gershon & A. S. '07
- Time dependent CP Bd $\Rightarrow K(\pi)\pi \gamma$; Bs $\Rightarrow \phi \gamma \sim O(10\%)$
- $\Delta \sin 2\beta(\text{penguins}) \sim O(\text{few } \%)$ i.e. comparable to QCD uncertainties.....
- $\Delta \gamma \sim O(2 \times 10^{-3})$ comparable to theory uncertainties

Null Tests of SM-CKM

Table 2. Illustrative sample of *approximate null tests (ANTs)*, with rough SM expectations and theoretical errors, current experimental uncertainties and estimates of numbers of B mesons needed for a Super B Factory to approach the SM uncertainty. More details for each mode can be found in the text.

Observable	SM expectation	Current experimental uncertainty ($B\bar{B}$ pairs used)	$B\bar{B}$ pairs needed
$\Delta S[\eta' K^0, \phi K^0, K^0 \bar{K}^0 K^0, \dots]$	$\sim(0 \pm 2)\%$	20% (6×10^8)	5×10^{10}
$\mathcal{A}_{CP}^{s+d}[M^0 X_{s+d}]$	$\lesssim 0.1\%$	—	$> 10^{12}$
$\mathcal{A}_{CP}[X_s \gamma]$	$(0.5 \pm 0.2)\%$	4% (2.4×10^8)	10^{11}
$\mathcal{A}_{CP}[X_d \gamma]$	$(-10 \pm 5)\%$	—	10^{11}
$\mathcal{A}_{CP}[X_{s+d} \gamma]$	$(0.000 \pm 0.001)\%$	12% (10^8)	$> 10^{12}$
$\mathcal{A}_{CP}[X_s l^+ l^-]$	$(-0.2 \pm 0.2)\%$	26% (10^8)	10^{12}
$\mathcal{A}_{CP}[X_d l^+ l^-]$	$(4 \pm 4)\%$	—	10^{12}
$\mathcal{A}_{CP}[X_{(s,d)} l^+ l^-]$	—	—	$> 10^{12}$
$\Sigma(\mathcal{A}_{CP}(\pi K))$	$(0 \pm 1)\%$	15% (6×10^8)	$> 10^{11}$
$\mathcal{A}_{CP}(\pi^+ \pi^0)$	$\lesssim 1\%$	6% (6×10^8)	10^{10}
$S[K_S \pi^0 \gamma, \dots]$	$\sim(0 \pm 5)\%$	28% (6×10^8)	$> 10^{10}$
$\langle p_t^{\tau} \rangle (D(X_c) \tau \nu_{\tau})$	0	—	$> 10^{12}$

ESP. relevant for
SKEK-B & SLHCb

T. Gershon + AS JHEP'07

(More) For The Intensity Frontier

- Charm CP esp. modes where SM predicts 0...e.g $D \Rightarrow KKX, \phi\pi^+, \pi^+\pi^0$
- ε'/ε : Hadronic matrix elements still a huge challenge
- $K \Rightarrow \pi^+ \nu \nu$; $KL \Rightarrow \pi^0 \nu \nu$

Desperate search for deviations from SM

For the Energy Frontier

- $t \Rightarrow c Z, ch \quad Br \sim O(10^{-7})$; $t \Rightarrow c g \sim O(10^{-5})$; $t \Rightarrow c \gamma \sim O(10^{-6})$many orders of magnitude bigger than SM
- $ee \Rightarrow tc$; $R_{tc} \sim 10^{-6} - 10^{-5}$
- $tedm \sim O(10^{-20} \text{ e-cm})$
- Triple correlation in $ee \Rightarrow tth$;
- $\Delta \text{ SM in } h \Rightarrow bb$

**Expected deviations tend
to be very small, strongly
suggesting we need to
strengthen both our
computational AND
measurement
infrastructure**

CP VIOLATION IN TOP PHYSICS

David ATWOOD^a, Shaouly BAR-SHALOM^b, Gad EILAM^c, Amarjit SONI^d

Containing studies of many reactions
for e^+e^- & pp e.g.
 $t\bar{t}dm$; $e^+e^- \rightarrow t\bar{c}$; $t\bar{t}h$...

Table 2: Real and Imaginary parts of d_t^γ and d_t^Z in units of 10^{-19} e-cm for $m_h = 100, 200$ and 300 GeV and for $\sqrt{s} = 500$ GeV and $\sqrt{s} = 1$ TeV (in parenthesis). $\tan \beta = 0.3$ and Set I,II,III means $\{\alpha_1, \alpha_2\} = \{\pi/4, \pi/2\}, \{\pi/4, 3\pi/4\}, \{\pi/4, \pi/4\}$, respectively.

Type of moment (10^{-19} e - cm) \downarrow	m_h (GeV) \downarrow	The different Sets of $\{\alpha_1, \alpha_2\}$, $\tan \beta = 0.3$		
		Set I	Set II	Set III
$\Re(d_t^\gamma)$	100	1.97(3.77)	1.40(2.66)	1.40(2.66)
	200	-3.36(2.26)	-2.38(1.60)	-2.38(1.60)
	300	-4.75(1.27)	-3.36(0.90)	-3.36(0.90)
$\Im(d_t^\gamma)$	100	-23.89(-5.44)	-16.88(-3.84)	-16.88(-3.84)
	200	-16.56(-4.91)	-11.70(-3.47)	-11.70(-3.47)
	300	-11.34(-4.33)	-8.02(-3.06)	-8.02(-3.06)
$\Re(d_t^Z)$	100	0.62(1.25)	0.36(0.83)	0.52(0.93)
	200	-1.17(0.74)	-0.87(0.47)	-0.78(0.57)
	300	-1.57(0.40)	-1.04(0.24)	-1.18(0.33)
$\Im(d_t^Z)$	100	-7.96(-1.81)	-5.41(-1.21)	-5.85(-1.34)
	200	-5.45(-1.62)	-3.58(-1.08)	-4.12(-1.22)
	300	-3.64(-1.42)	-2.22(-0.93)	-2.91(-1.08)

See also
Gunion,
Grazdowski
& He, PRL 96

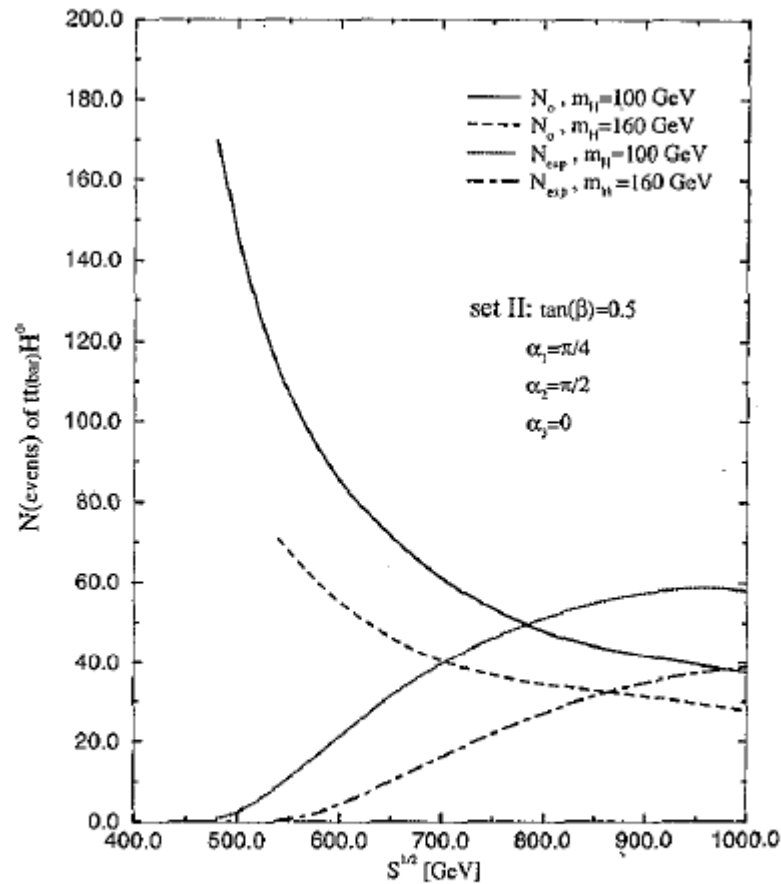


FIG. 3. Number of events, N_O (N_{exp}) required (expected yearly), as a function of total beam energy for set II of the parameters and for $m_{H^0} = 100$ and 160 GeV with unpolarized electron and positron beams.

$e^+e^- \rightarrow t\bar{t}h$
Triple
Correlations

Atwood et al
PRD 96

Perhaps feasible at $\sqrt{s} \sim 750$ GeV ILC

**Constraining the FC $htc(u)$ vertex
in the context of RS [i.e. (partial)
compositeness] is of special
significance**

Constraining FCNC: Exploit (PARTIAL) COMPOSITENESS

Agashe et al'06

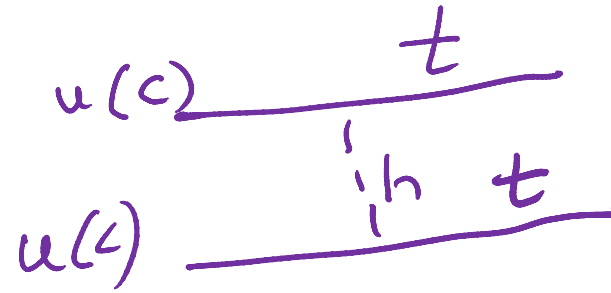
CSAKI ET AL JHEP'08

AGASHE ET AL PRD'07

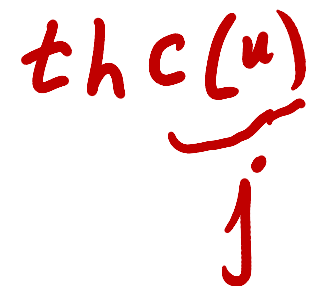
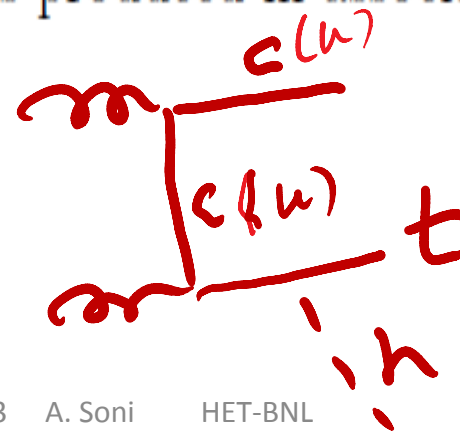
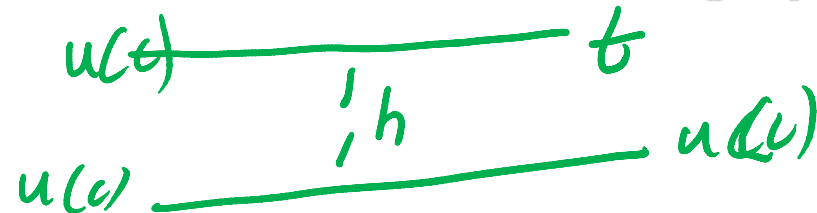
AGASHE, CONTINO, PRD'09

AZATOV ET AL PRD'09

KEREN-ZUR ET AL, NPB '13



- (a) the same-sign top-pair, e.g. $pp \rightarrow tt(\bar{t}\bar{t})$,
- (b) processes where a top-quark is produced in association with a light jet, e.g. $pp \rightarrow t\bar{j}_u(\bar{t}j_u)$, where $j_u = u, c$, and,
- (c) processes where the top-quark is produced in association with a Higgs and a light jet, e.g. $pp \rightarrow t\bar{j}_u h(\bar{t}j_u h)$.



Effective tree-level

$$\sqrt{\xi_{tc}^2 + \xi_{tu}^2} \lesssim \begin{cases} 0.25 \text{ (0.2)} & \text{for process (a)} \\ 0.9 \text{ (0.9)} & \text{for process (b)} \\ 0.26 \text{ (0.1)} \times 10^{-4} & \text{for process (c)} \end{cases}$$

$pp \rightarrow tqh$

$\sqrt{s} = 7 \text{ TeV},$
5/fb

$\sqrt{s} = 8 \text{ TeV},$ 22/fb

Production level constraints

[RS bias $\{t_c\} \gg \{t_u\}$]

$$\sqrt{\xi_{tc}^2 + \xi_{tu}^2} \lesssim \begin{cases} 0.47 \text{ (0.36)} & \text{for } pp \rightarrow l^\pm l^\pm + 2b - jets + \cancel{E}_T \\ 0.9 \text{ (0.9)} & \text{for } pp \rightarrow l^\pm + j + b - jet + \cancel{E}_T \\ 4.5 \text{ (1.8)} \times 10^{-4} & \text{for } pp \rightarrow l^\pm + j + 3b - jets + \cancel{E}_T \text{ * } \\ 52.5 \text{ (19.8)} \times 10^{-4} & \text{for } pp \rightarrow l^\pm + j + b - jets + 2\gamma + \cancel{E}_T \end{cases} \quad (5.2)$$

* $\sqrt{s} = 7 \text{ TeV}, 5/\text{fb}$

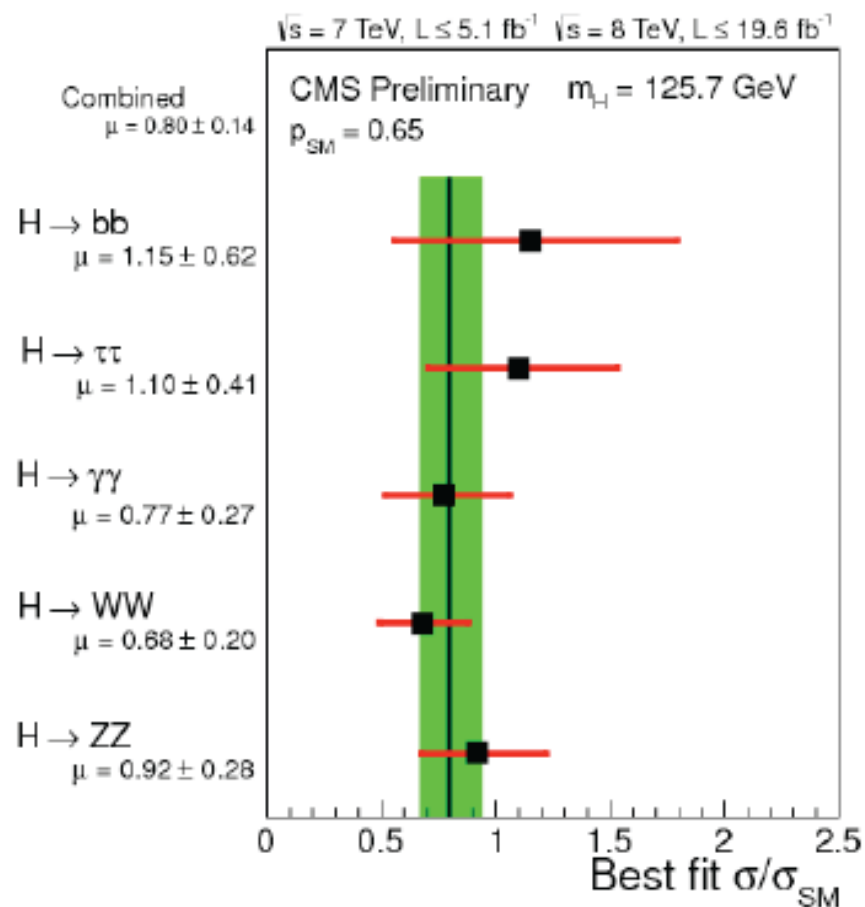
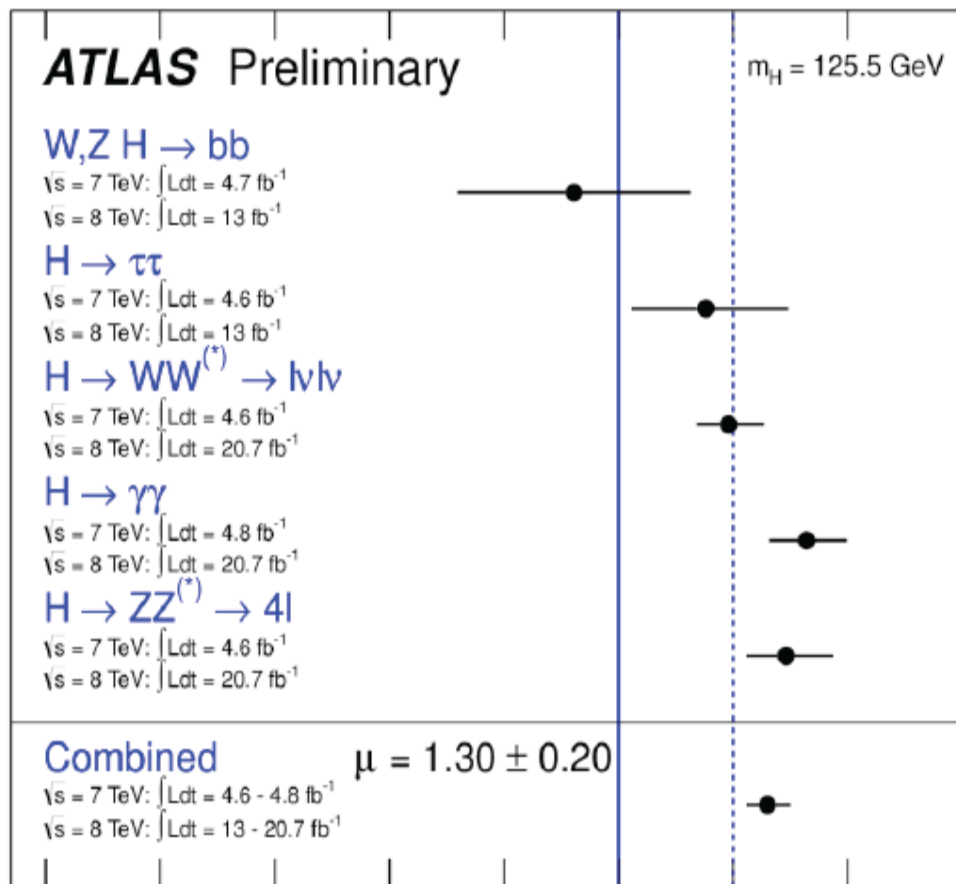
$(\sqrt{s} = 8 \text{ TeV}, 22/\text{fb})$

Harnik, Kopp, Zupan
arXiv:1209.1397

Technique	Coupling	Constraint
D^0 oscillations [48]	$ Y_{uc} ^2, Y_{cu} ^2$	$< 5.0 \times 10^{-9}$
	$ Y_{uc}Y_{cu} $	$< 7.5 \times 10^{-10}$
B_d^0 oscillations [48]	$ Y_{db} ^2, Y_{bd} ^2$	$< 2.3 \times 10^{-8}$
	$ Y_{db}Y_{bd} $	$< 3.3 \times 10^{-9}$
B_s^0 oscillations [48]	$ Y_{sb} ^2, Y_{bs} ^2$	$< 1.8 \times 10^{-6}$
	$ Y_{sb}Y_{bs} $	$< 2.5 \times 10^{-7}$
K^0 oscillations [48]	$\text{Re}(Y_{ds}^2), \text{Re}(Y_{sd}^2)$	$[-5.9 \dots 5.6] \times 10^{-10}$
	$\text{Im}(Y_{ds}^2), \text{Im}(Y_{sd}^2)$	$[-2.9 \dots 1.6] \times 10^{-12}$
	$\text{Re}(Y_{ds}^* Y_{sd})$	$[-5.6 \dots 5.6] \times 10^{-11}$
	$\text{Im}(Y_{ds}^* Y_{sd})$	$[-1.4 \dots 2.8] \times 10^{-13}$
single-top production [49]	$\sqrt{ Y_{tc}^2 + Y_{ct} ^2}$	< 3.7
	$\sqrt{ Y_{tu}^2 + Y_{ut} ^2}$	< 1.6
$t \rightarrow hj$ [50]	$\sqrt{ Y_{tc}^2 + Y_{ct} ^2}$	< 0.34
	$\sqrt{ Y_{tu}^2 + Y_{ut} ^2}$	< 0.34
D^0 oscillations [48]	$ Y_{ut}Y_{ct} , Y_{tu}Y_{tc} $	$< 7.6 \times 10^{-3}$
	$ Y_{tu}Y_{ct} , Y_{ut}Y_{tc} $	$< 2.2 \times 10^{-3}$
	$ Y_{ut}Y_{tu}Y_{ct}Y_{tc} ^{1/2}$	$< 0.9 \times 10^{-3}$
neutron EDM [37]	$\text{Im}(Y_{ut}Y_{tu})$	$< 4.4 \times 10^{-8}$

) $pp \rightarrow t h j$ gives
much tighter
constraints.

Table 2. Constraints on flavor violating Higgs couplings to quarks. We have assumed a Higgs mass $m_h = 125 \text{ GeV}$, and we have taken the diagonal Yukawa couplings at their SM values.



$\mu = 0.80 \pm 0.14$

E. LIPELES
PBF2013

S. BOSE PBF2013

Is Nature Unnatural?

Decades of confounding experiments have physicists considering a startling possibility: The universe might not make sense.

by: [Natalie Wolchover](#)

May 24, 2013

[email](#) [print](#)



Is the universe natural or do we live in an atypical bubble in a multiverse? Recent results at the Large Hadron Collider have forced many physicists to confront the latter possibility. (Illustration: Giovanni Villadoro)

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**Gee, don't see no NP signals
Flavor: Told you so!**

KEY MESSAGES FROM A CANDIDATE THEORY OF FLAVOR

Key messages from a candidate theory of flavor

1. In a candidate theory, the gigantic tension between hierarchy and flavor puzzle gets dramatically ameliorated. *Thus remarkably RS-leads to lowering of Λ_{flavor} from ~ 1000 to ~ 10 TeV*

Beat them to Death!

II. Due to flavor mis-alignment, $O(1)$ BSM phases occur naturally; \Rightarrow direct CP is an extremely powerful probe of flavor alignment and holds the key to unlocking new physics. For this purpose, fortunately, there are many observables : $\text{Re}(\epsilon'/\epsilon)$; γ ; $S[B \Rightarrow K^*(\rho)\gamma]$; $\Delta \sin 2\beta$ from $B_d \Rightarrow \eta' K_s$, ϕK_s , $3 K_s$; $\text{ACP}(B \Rightarrow K\pi)$, $\Delta \text{ACP}(D)$but expected signals tend to be small necessitating high precision.

Beat Them to Death!

- III Top quark edm may be non-vanishing and its measurement deserves special attention
- IV Top quark is very sensitive to flavor violation;
 $t \Rightarrow c Z$; $t \Rightarrow c h$, $pp \Rightarrow t c h X$ etc need to be vigorously pursued.
- V. Lepton flavor violation is a natural prediction \Rightarrow
Searches for $\tau = \mu \gamma$, $3 \mu \dots$; $B_s \Rightarrow \tau \mu \dots$ are very important. [See Agashe, Blechman & Petriello'06]

VI. Expected size of corrections to Higgs couplings

- Deviation from SM $\sim O(v^2 / m_{KK}^2) \sim 0.5\%$
[assuming $m_{KK} > \sim 10 \text{ TeV}$].

Such small corrections should be a concern.

- VII. For direct observation of KK-particles of mass $> \sim 10 \text{ TeV}$ need a Gigantic International Hadron Collider (GIHC) $\sim 100 \text{ TeV}$ cm energy

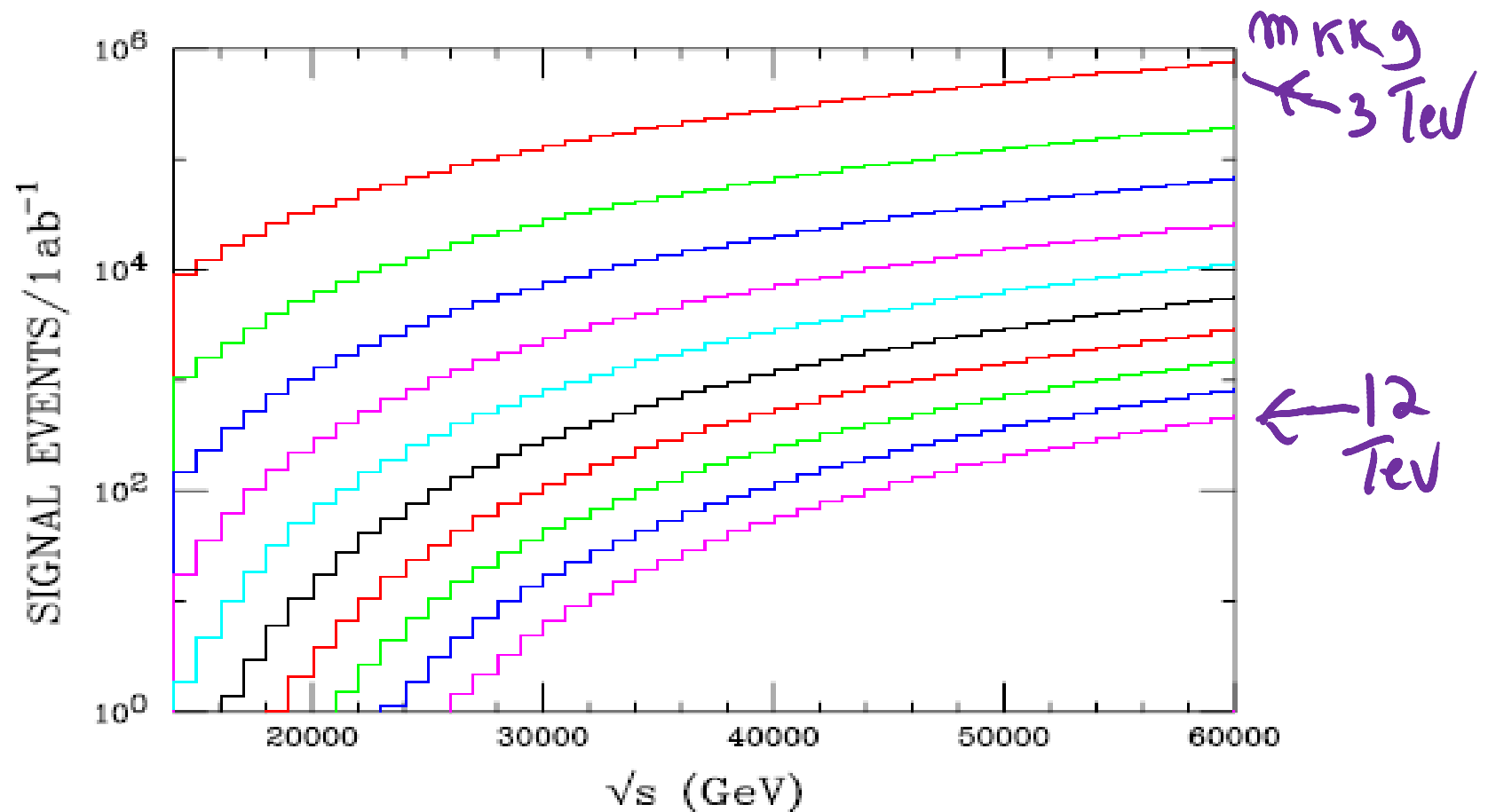


FIG. 10 (color online). Signal rate for a possible gluon KK resonance as a function of the collider energy employing the cuts described in the text. Branching fractions and efficiencies have been neglected. From top to bottom, the results are shown for gluon KK masses in the range from 3 to 12 TeV in steps of 1 TeV.

Lesson learnt from ν 's

~ Circa 1983, after long and arduous efforts, Δm^2 upper bound used to be around a few eV^2 but efforts to search oscillations continued basically because there was no good theoretical reason for m_ν to be zero.

- *Recall it took more than a decade beyond '83* and Δm^2 had to be lowered by almost 4 orders of magnitude (!) before osc were discovered.
- **Moral: Physical “principles” [i.e Naturalness in this instance] shouldn't be abandoned easilyWe'll just have to work harder to get to it but this should have been anticipated if enough attention had been paid to flavor alignment**

Recall SSC ~ 40 TeV 1990 technologically
completely feasible.

We should be SERIOUSLY

THINKING of

GIGANTIC INTERNATIONAL
HADRON COLLIDER [GIHC]
 ~ 100 TeV CM

↓
"GEEK"

Summary & Outlook (p1 of 2)

- While naturalness is not tangible, [clearly 10^{-2} OR 10^{-4} are very different from 10^{-34}], flavor-alignment places specific constraints...has been telling us for long that scale of NP > 1 TeV
- Specifically RS-flavor (which gives a nice geometric understanding of flavor & simultaneously of EW-Plank hierarchy) strongly suggests scale is unlikely less than ~ 10 TeV.
- At ~ 10 TeV tuning is worse, but (as a bonus) the theory is simpler; EWP are automatically satisfied.

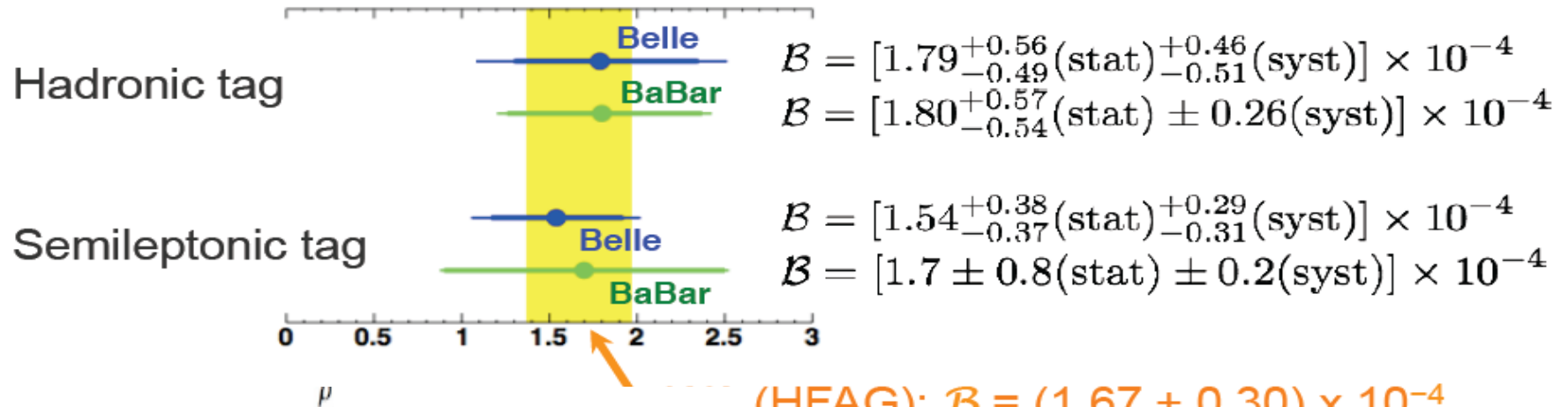
Summary & Outlook (p.2)

- From the perspective of such a theory the following deserve attention
- Dir CP probes [e.g. nedm , ε'/ε , $S[B \Rightarrow K \rho \gamma]$; γ ; Null Tests, D-CP...
- t-dm; top FV via e.g. $t \Rightarrow c Z$; $t \Rightarrow c h$; $pp \Rightarrow t c h$; $e^+ e^- \Rightarrow t c$
- τ FV: $\tau \Rightarrow \mu \gamma$; 3μ ; $B_s \Rightarrow \tau \mu$
- Expected deviation to higgs couplings $\sim O(0.5\%)$ should be a concern in planning $e^+ e^-$ collider.
- Precise measurements & precise computations deserve high priority.
- It is essential to have high sensitivity flavor experiments AND we should be seriously thinking of a GILHC as the next step in our adventure.

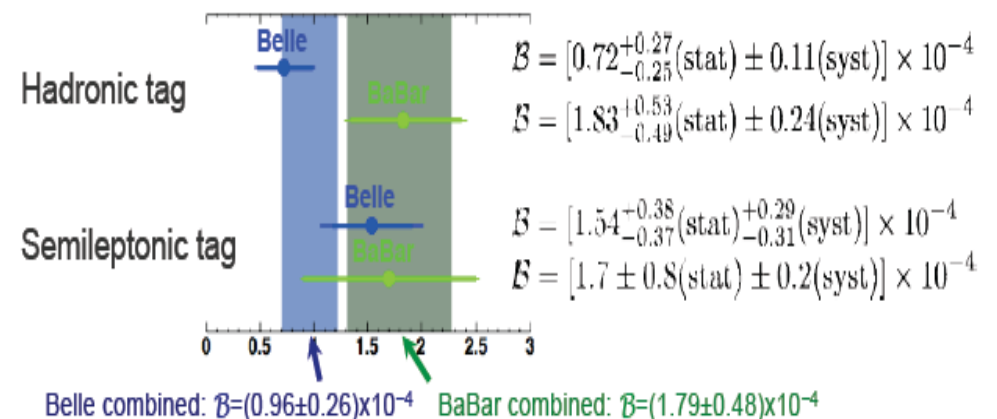
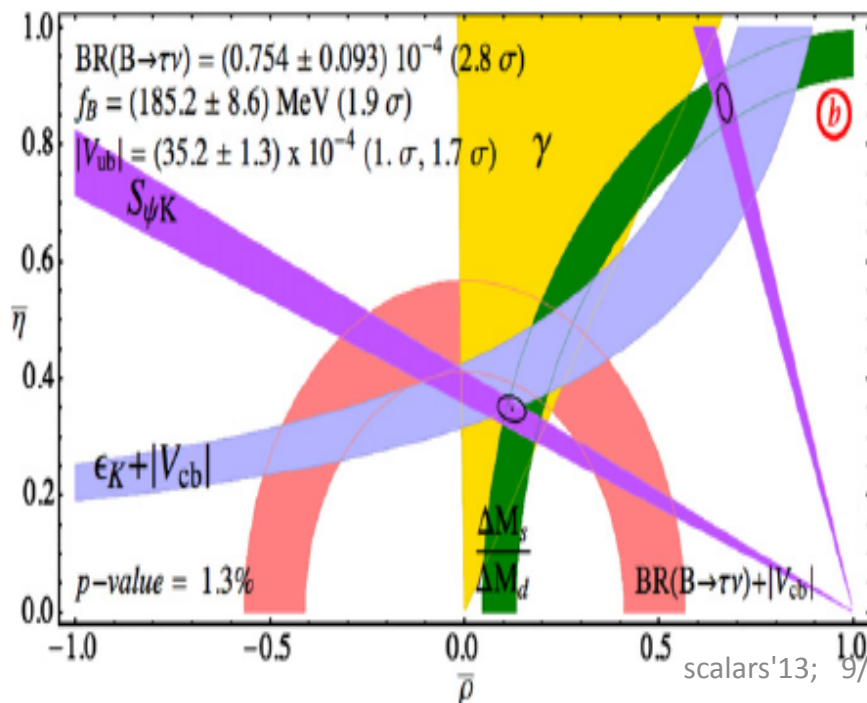
XTRA

ANOMALIES (A SAMPLE)

Status for $B \rightarrow \tau \nu$ before ICHEP 2012



Status for $B \rightarrow \tau \nu$ after ICHEP 2012



A naive world average: $\mathcal{B} = (1.15 \pm 0.23) \times 10^{-4}$

24 Nov 11 + APS MLB '11

$B \rightarrow D^{(*)} \tau \nu$ from BaBar and SM

$$\mathcal{R}(D)_{\text{exp}} = 0.440 \pm 0.072 \quad \mathcal{R}(D^*)_{\text{exp}} = 0.332 \pm 0.030$$

$$\updownarrow 2.0\sigma$$

$$\updownarrow 2.7\sigma$$

$$\mathcal{R}(D)_{\text{SM}} = 0.297 \pm 0.017 \quad \mathcal{R}(D^*)_{\text{SM}} = 0.252 \pm 0.003$$

SM expectations in S. Fajfer, J. Kamenik, I. Nisandzic, PRD 85, 094025 (2012).

$B \rightarrow D^{(*)} \tau \nu$ from Belle

A. Bozek's averages (KEK-FF 2013):

(naive averages for inclusive and exclusive hadronic tags)

$$R(D) = 0.430 \pm 0.091$$

Deviation from SM

$$1.4 \sigma$$

$$R(D^*) = 0.405 \pm 0.047$$

$$3.0 \sigma$$

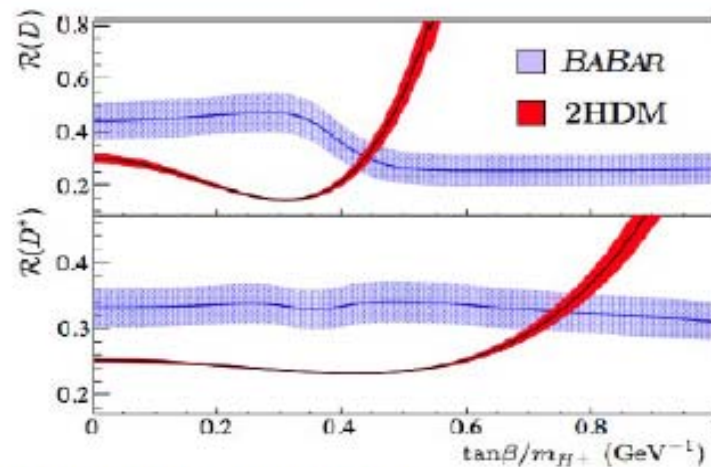
$$\text{Combined } 3.3 \sigma$$

Correlation btw $R(D)$ and $R(D^*)$ neglected conservatively.

[B BHUYAN]



MSSM: 2HDM-Type II



PRD 78, 015006 (2008)
PRD 85, 094025 (2012)

$$\tan\beta/m_{H^+} = 0.44 \pm 0.02 \text{ GeV}^{-1}$$

$$\tan\beta/m_{H^+} = 0.75 \pm 0.04 \text{ GeV}^{-1}$$

$R(D)$ and $R(D^*)$ do not both agree with the predictions for any single value of $\tan\beta/m_{H^+}$

2HDM-Type-II excluded over the full parameter space* at 99.8% CL.

* relies on $m_{H^+} < 15 \text{ GeV}$ being excluded based on $b \rightarrow s\gamma$

VIIIth Rencontres du Vietnam

Bipul Bhuyan, IIT Guwahati

17

Perhaps premature

If true another huge stat at SUSY

FORWARD-BACKWARD TOP ASY

SM e NNLO⁺
 0.072 ± 0.011
 -0.007

N2

Observable	Values	Experiment
\mathcal{A}_{FB}^t	0.19 ± 0.065	DØ Collaboration [1]
	0.158 ± 0.074	CDF Collaboration [2]
	0.176 ± 0.05	Combined
$\mathcal{A}_{FB}^{t,low}$	0.078 ± 0.048	DØ Collaboration [1]
	-0.022 ± 0.043	CDF Collaboration [2]
	0.023 ± 0.032	Combined
$\mathcal{A}_{FB}^{t,high}$	0.115 ± 0.060	DØ Collaboration [1]
	0.266 ± 0.062	CDF Collaboration [2]
	0.188 ± 0.043	Combined
$\sigma_{t\bar{t}}^{Tevatron}$	$8.18^{+0.98}_{-0.87}$ pb	DØ Collaboration [8]
$\sigma_{l\pm l\pm}^{LHC}$	< 1 fb	ATLAS & CMS Collaborations [9]

Table 1: Measured values of various observables used in our analysis; combined here mean weighted averages.

+

V. Ahrens, A. Ferroglia, M. Neubert, B. D. Pecjak and L. L. Yang, JHEP **1009**, 097 (2010) [arxiv:1003.5827 [hep-ph]].

Atwood, Gupta, AS, arXiv1301.2250

Berger et al 1101.5625

Aguilar-Saavedra, Perez-Victoria, 1104.1385

Degrande et al 1104.1798 56

powerful discriminator SST

scalars'13; 9/13/13 A. So