

Higgs decays to $\gamma\gamma$ and $Z\gamma$ in the Inert Doublet Model

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Phys. Rev. D 88, 035019 (2013), JHEP 09 (2013) 055

Outline

- Motivation
- Introduction to IDM
- $h \rightarrow \gamma\gamma, Z\gamma$ rates and their correlation
- $R_{\gamma\gamma} > 1$
 - Bounds on masses
- $1 > R_{\gamma\gamma} > 0.7 \rightarrow$ see the talk of D. Sokołowska
 - Implications for the DM
 - Combination with WMAP
- Summary

Why $h \rightarrow \gamma\gamma$?

- Important Higgs decay channel at the LHC
- Experimental hints on deviation from the SM value

ATLAS : $R_{\gamma\gamma} = 1.65 \pm 0.24(\text{stat})^{+0.25}_{-0.18}(\text{syst})$

CMS : $R_{\gamma\gamma} = 0.79^{+0.28}_{-0.26}$

- Sensitivity to new charged particles – well suited for studying different 2HDMs
- Sensitivity to invisible decay channels – information about extra states

→ also $h \rightarrow Z\gamma$ interesting, but not enough data

Why Inert Doublet Model?

- Rich phenomenology
- $\rho = 1$ at the tree-level
- Viable DM candidate
- Thermal evolution of the Universe
+ some conditions for baryogenesis

Inert Doublet Model

[N. G. Deshpande, E. Ma, Phys. Rev. D 18 (1978) 2574, J. F. Gunion, H. E. Haber, G. Kane, S. Dawson, *The Higgs Hunter's Guide*, 1990 Addison-Wesley, R. Barbieri, L. J. Hall, V. S. Rychkov, Phys. Rev. D74 (2006) 015007, I. F. Ginzburg, K. A. Kanishev, M. Krawczyk, D. Sokolowska, Phys. Rev. D 82 (2010) 123533]

IDM – a 2HDM with the scalar potential (real parameters):

$$V = -\frac{1}{2} \left[m_{11}^2 (\phi_S^\dagger \phi_S) + m_{22}^2 (\phi_D^\dagger \phi_D) \right] + \frac{1}{2} \left[\lambda_1 (\phi_S^\dagger \phi_S)^2 + \lambda_2 (\phi_D^\dagger \phi_D)^2 \right] + \\ + \lambda_3 (\phi_S^\dagger \phi_S)(\phi_D^\dagger \phi_D) + \lambda_4 (\phi_S^\dagger \phi_D)(\phi_D^\dagger \phi_S) + \frac{1}{2} \lambda_5 \left[(\phi_S^\dagger \phi_D)^2 + (\phi_D^\dagger \phi_S)^2 \right]$$

- **D symmetry:** $\phi_D \rightarrow -\phi_D, \phi_S \rightarrow \phi_S$
- Yukawa interactions: type I (only ϕ_S couples to fermions)
- \mathcal{L} – D-symmetric
- **D-symmetric vacuum state** $\langle \phi_S \rangle = \frac{v}{\sqrt{2}}, \langle \phi_D \rangle = 0$

⇒ EXACT D-symmetry

- Number of theoretical and experimental constraints

Particle spectrum of IDM

[E. M. Dolle, S. Su, Phys. Rev. D 80 (2009) 055012, L. Lopez Honorez, E. Nezri, F. J. Oliver, M. Tytgat, JCAP 0702 (2007) 028, D. Sokołowska, arXiv:1107.1991 [hep-ph]]

- ϕ_S : h – SM-like Higgs boson, tree-level couplings to fermions and gauge bosons like in the SM.
Deviation from SM in loop couplings possible!
- ϕ_D : H, A, H^\pm – dark scalars, no tree-level couplings to fermions
- D symmetry **exact** \Rightarrow lightest D -odd particle stable
 \Rightarrow **DM candidate**
- DM = H , so $M_H < M_{H^\pm}, M_A$
- Three regions of DM mass consistent with astrophysical observations (WMAP: $0.1018 < \Omega_{DM} h^2 < 0.1234$):
 - $M_H \lesssim 10$ GeV
 - 40 GeV $< M_H < 150$ GeV
 - $M_H \gtrsim 500$ GeV

$\gamma\gamma$ and $Z\gamma$ decay rates of the Higgs boson

[Q.-H. Cao, E. Ma, G. Rajasekaran, Phys. Rev. D 76 (2007) 095011, P. Posch, Phys. Lett. B 696 (2011) 447, A. Arhrib, R. Benbrik, N. Gaur, Phys. Rev. D 85 (2012) 095021, BŚ. M. Krawczyk, Phys. Rev. D 88 (2013) 035019]

$R_{\gamma\gamma}$ – 2-photon decay rate, $R_{Z\gamma}$ – $Z\gamma$ decay rate

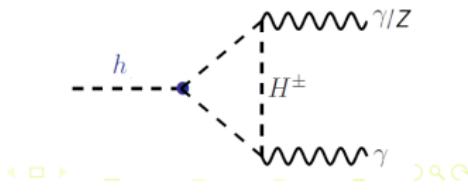
$$R_{\gamma\gamma} = \frac{\sigma(pp \rightarrow h \rightarrow \gamma\gamma)^{IDM}}{\sigma(pp \rightarrow h \rightarrow \gamma\gamma)^{SM}} \approx \frac{\Gamma(h \rightarrow \gamma\gamma)^{IDM}}{\Gamma(h \rightarrow \gamma\gamma)^{SM}} \frac{\Gamma(h)^{SM}}{\Gamma(h)^{IDM}}$$

$R_{Z\gamma}$ – treated analogously

- Largest contribution from gg fusion
 - $\sigma(gg \rightarrow h)^{SM} = \sigma(gg \rightarrow h)^{IDM}$ (not true in other 2HDMs)

Two sources of deviation from $R_{yy} = 1$:

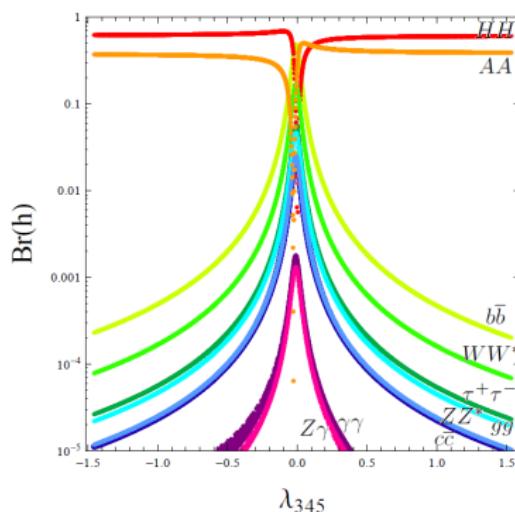
- **invisible decays** $h \rightarrow HH, h \rightarrow AA$
in $\Gamma(h)^{IDM}$
 - **charged scalar loop** in $\Gamma(h \rightarrow \gamma\gamma)^{IDM}$



Invisible decays

$$\begin{aligned}\Gamma(h) = & \Gamma(h \rightarrow b\bar{b}) + \Gamma(h \rightarrow WW^*) + \Gamma(h \rightarrow \tau^+\tau^-) + \Gamma(h \rightarrow gg) \\ & + \Gamma(h \rightarrow ZZ^*) + \Gamma(h \rightarrow c\bar{c}) + \Gamma(h \rightarrow Z\gamma) + \Gamma(h \rightarrow \gamma\gamma) \\ & + \textcolor{red}{\Gamma(h \rightarrow HH)} + \textcolor{blue}{\Gamma(h \rightarrow AA)}\end{aligned}$$

- Controlled by: M_H , M_A ,
 $\lambda_{345} \sim hHH$, $\lambda_{345}^- \sim hAA$
 - Invisible decays, if kinematically allowed, dominate over SM channels.
 - Plot for $M_A = 58$ GeV,
 $M_H = 50$ GeV

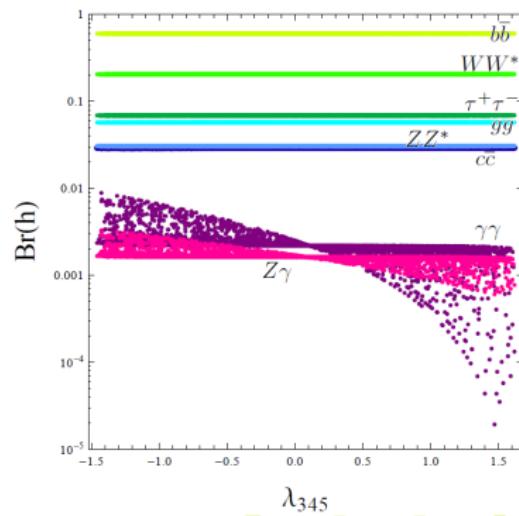


Charged scalar H^\pm loop

[J. R. Ellis, M. K. Gaillard and D. V. Nanopoulos, Nucl. Phys. B 106 (1976) 292, M. A. Shifman, A. I. Vainshtein, M. B. Voloshin and V. I. Zakharov, Sov. J. Nucl. Phys. 30 (1979) 711 [Yad. Fiz. 30, 1368 (1979)]]

$$\Gamma(h \rightarrow \gamma\gamma)^{IDM} = \frac{G_F \alpha^2 M_h^3}{128\sqrt{2}\pi^3} \left| \mathcal{A}^{SM} + \frac{2M_{H^\pm}^2 + m_{22}^2}{2M_{H^\pm}^2} A_0 \left(\frac{4M_{H^\pm}^2}{M_h^2} \right) \right|^2$$

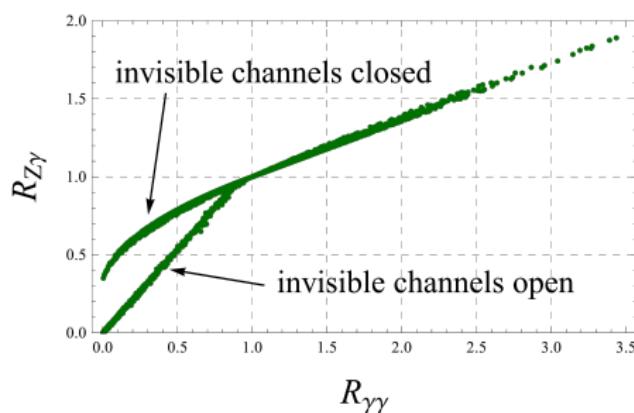
- Constructive or destructive interference between SM and H^\pm contributions
- Controlled by M_{H^\pm} and $2M_{H^\pm}^2 + m_{22}^2 \sim \lambda_3 \sim hH^+H^-$
- **Invisible channels closed $\Rightarrow H^\pm$ contribution visible**



$h \rightarrow \gamma\gamma$ vs $h \rightarrow Z\gamma$

[BŚ, M. Krawczyk, Phys. Rev. D 88 (2013) 035019, formulas for $h \rightarrow Z\gamma$: A. Djouadi, Phys.Rept. 459 (2008) 1, C.-S. Chen, C.-Q. Geng, D. Huang, L.-H. Tsai, Phys.Rev.D 87 (2013) 075019]

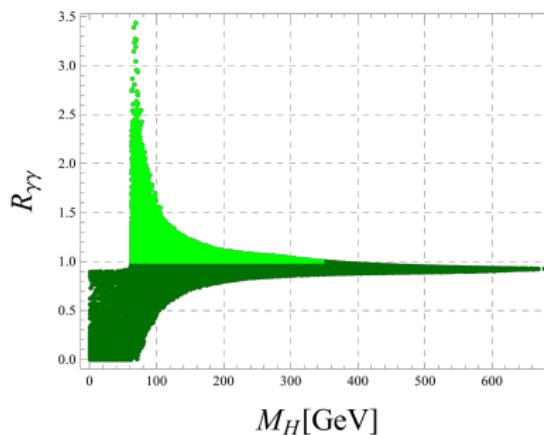
- Sensitivity to invisible channels
- $R_{\gamma\gamma}$ and $R_{Z\gamma}$ positively correlated
- $R_{\gamma\gamma} > 1 \Leftrightarrow R_{Z\gamma} > 1$



$R_{\gamma\gamma}$ vs Dark Matter (H) mass

[A. Arhrib, R. Benbrik, N. Gaur, Phys. Rev. D 85 (2012) 095021, BŚ, M. Krawczyk, Phys. Rev. D 88 (2013) 035019]

- Invisible channels open \Rightarrow
no enhancement in
 $h \rightarrow \gamma\gamma$ possible
- Enhanced $R_{\gamma\gamma}$ for
 $M_H, M_{H^\pm}, M_A > 62.5$ GeV
- $R_{\gamma\gamma} > 1 \Rightarrow$ very light DM
excluded

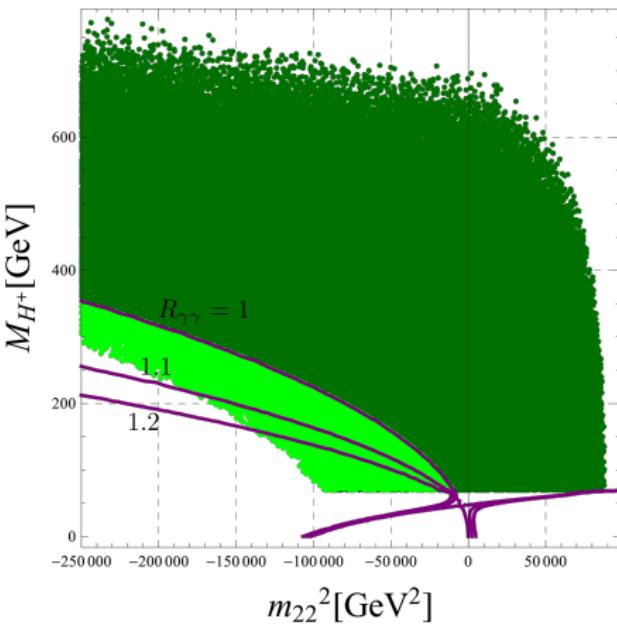


$R_{\gamma\gamma}$ vs charged scalar mass

Enhanced $R_{\gamma\gamma}$ even for big values of M_{H^\pm}

$$R_{\gamma\gamma} > 1.2 \Rightarrow M_{H^\pm}, M_H \lesssim 154 \text{ GeV}$$

- Only medium DM mass
- Light charged scalar

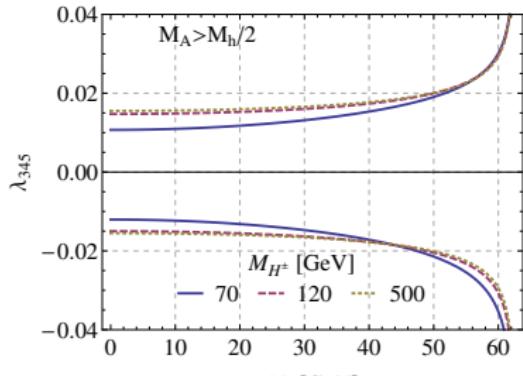
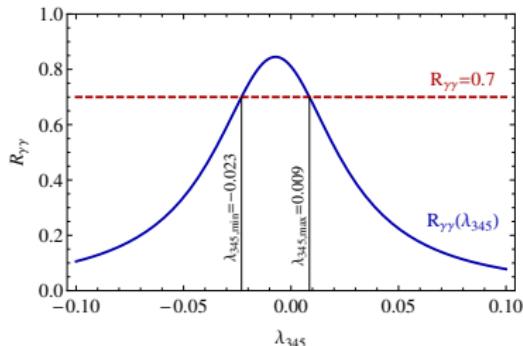


Invisible channel $h \rightarrow HH$ open

$R_{\gamma\gamma}$ constraints on $\lambda_{345} \sim hHH$

[M. Krawczyk, D. Sokołowska, P. Swaczyna, BŚ, JHEP 09 (2013) 055]

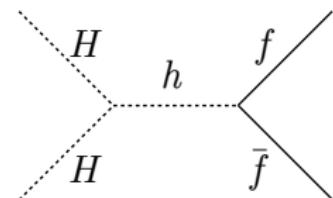
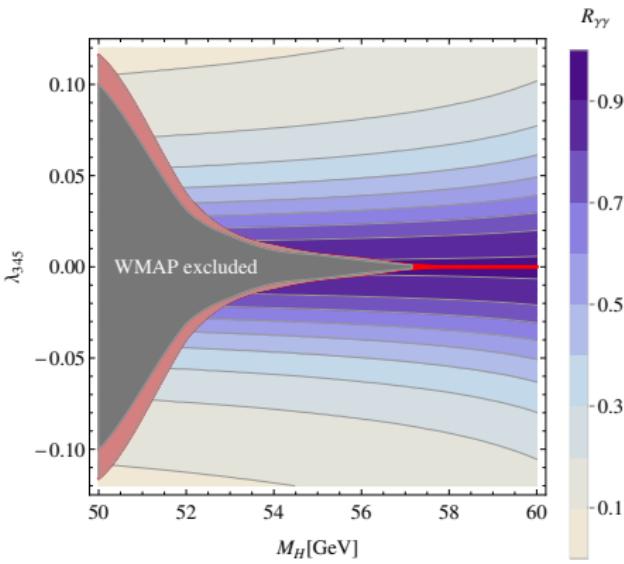
- Setting a lower limit on $R_{\gamma\gamma}$ constrains λ_{345}
 - Upper and lower limits on λ_{345} depend on M_H
 - Stronger than limits on $\text{Br}(h \rightarrow \text{inv})$ from LHC
 - Stronger than limits from XENON100



Invisible channel $h \rightarrow HH$ open

WMAP constraints on λ_{345}

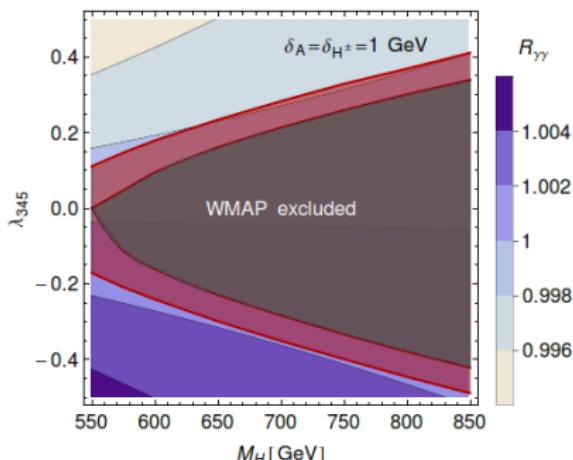
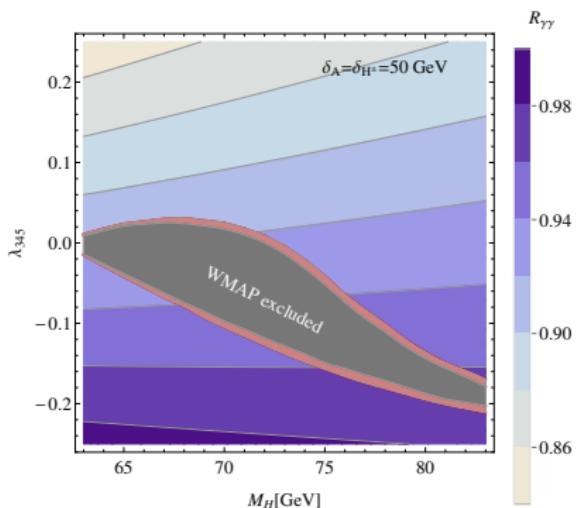
λ_{345} controls the annihilation of the DM
 $HH \rightarrow h \rightarrow f\bar{f} \Rightarrow$ important for the relic density of the DM



- $R_{\gamma\gamma} > 0.7$ inconsistent with WMAP for $M_H < 53$ GeV
- **Light DM excluded!**

Invisible channels closed

Intermediate and heavy DM



- H of intermediate mass can constitute 100% of DM
- H constituting 100% DM inconsistent with $R_{\gamma\gamma} > 1$

- For heavy DM $R_{\gamma\gamma} \approx 1$ only very small deviations allowed

Summary

- IDM in agreement with the data (LEP, LHC and WMAP)
- $R_{\gamma\gamma}$ and $R_{Z\gamma}$ positively correlated, $R_{\gamma\gamma} > 1 \Leftrightarrow R_{Z\gamma} > 1$
- $h \rightarrow \gamma\gamma$ can provide important information about IDM, because it is sensitive to M_H and M_{H^\pm}
- If substantial enhancement of $R_{\gamma\gamma}$
 - Only medium masses of DM
 - Light charged scalar
 - H – subdominant component of the DM
- If H constitutes 100% of DM
 - Light DM excluded
 - Intermediate DM can accommodate only $R_{\gamma\gamma} < 1$
 - For heavy DM $R_{\gamma\gamma} \approx 1$

Motivation

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IDM

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$h \rightarrow \gamma\gamma, Z\gamma$ rates

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$R_{\gamma\gamma} > 1$

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$1 > R_{\gamma\gamma} > 0.7$

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Summary

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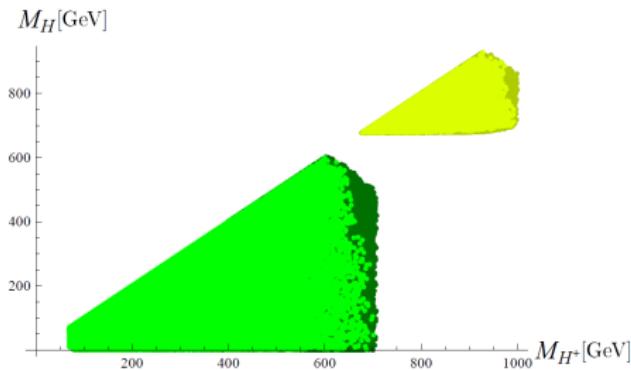
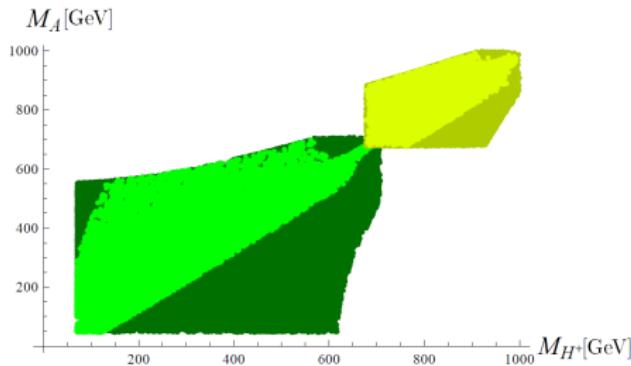
Back up

Constraints

- **Vacuum stability:** scalar potential V bounded from below
- **Perturbative unitarity:** eigenvalues Λ_i of the high-energy scattering matrix fulfill the condition $|\Lambda_i| < 8\pi$
- **Existence of the Inert vacuum:** Inert state – a global minimum of the scalar potential $\Rightarrow m_{22}^2 \lesssim 9 \cdot 10^4 \text{ GeV}^2$
- **H as DM candidate:** $M_H < M_A, M_{H^\pm}$ and WMAP
- **Electroweak Precision Tests (EWPT):** S and T within 2σ ($S = 0.03 \pm 0.09$, $T = 0.07 \pm 0.08$, with correlation of 87%)
- **LEP bounds** on the scalars' masses
- **LHC:** $M_h \approx 125 \text{ GeV}$

Masses of the scalars

$$\begin{aligned} M_h^2 &= m_{11}^2 = \lambda_1 v^2 \\ M_{H^\pm}^2 &= \frac{1}{2}(\lambda_3 v^2 - m_{22}^2) \\ M_A^2 &= \frac{1}{2}(\lambda_{345}^- v^2 - m_{22}^2) \\ M_H^2 &= \frac{1}{2}(\lambda_{345} v^2 - m_{22}^2) \end{aligned}$$



$R_{\gamma\gamma} > 1$ – analytical solution

If invisible channels closed

$$R_{\gamma\gamma} = \frac{\Gamma(h \rightarrow \gamma\gamma)^{\text{IDM}}}{\Gamma(h \rightarrow \gamma\gamma)^{\text{SM}}}$$

$\Rightarrow R_{\gamma\gamma} > 1$ can be solved analytically for M_{H^\pm} , m_{22}^2

• Constructive interference

- $m_{22}^2 < -2M_{H^\pm}^2$ ($\Leftrightarrow \lambda_3 < 0$)
- with LEP bound on $M_{H^\pm} \Rightarrow$
 $m_{22}^2 < -9.8 \cdot 10^3 \text{ GeV}^2$

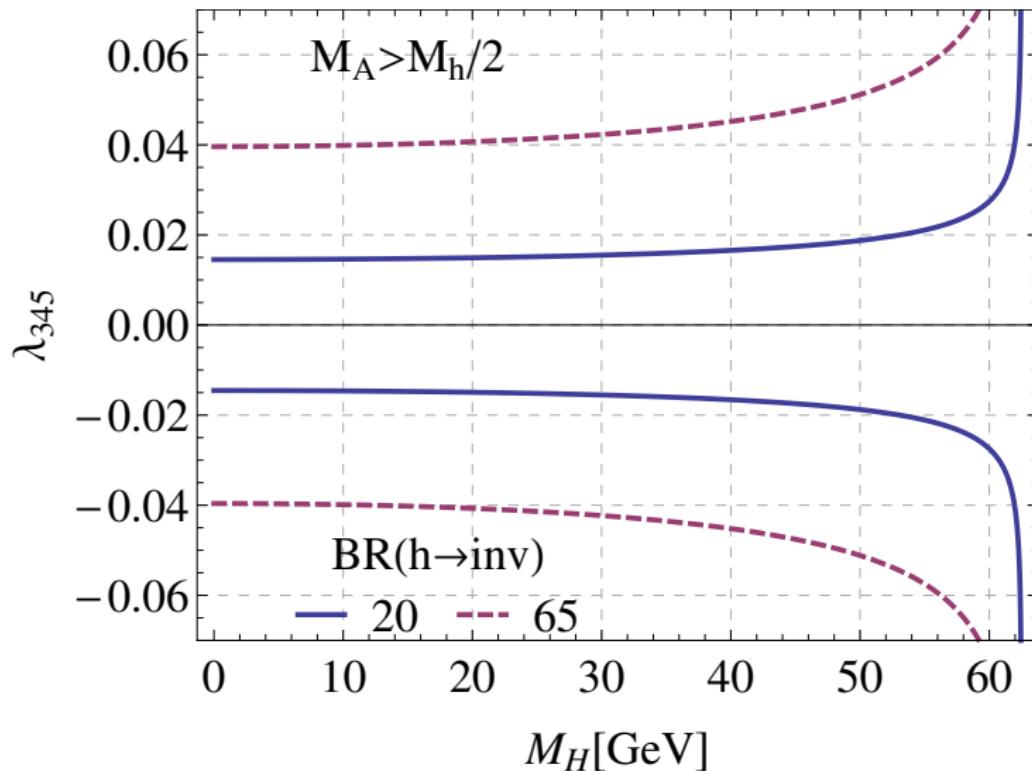
• Destructive interference

- IDM contribution $\geq 2 \times$ SM contribution
- big m_{22}^2 required:
 $m_{22}^2 \gtrsim 1.8 \cdot 10^5 \text{ GeV}^2$
- excluded** by the condition for the Inert vacuum
 $m_{22}^2 \lesssim 9 \cdot 10^4 \text{ GeV}^2$

DM signals

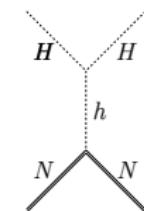
[see e.g.: M. Gustafsson, S. Rydbeck, L. Lopez Honorez, E. Löndstrom, Phys. Rev. D 86 (2012) 075019]

- gamma-ray lines
- cosmic and neutrino fluxes
- direct detection signals

Limits on λ_{345} from $\text{Br}(h \rightarrow \text{inv})$ 

Comparison with XENON

DM-nucleon scattering cross section $\sigma_{\text{DM},N} \sim \lambda_{345}^2$



- $R_{\gamma\gamma}$ bounds on λ_{345} translated to the $(\sigma_{\text{DM},N}, M_H)$ plane
- Upper limits are stronger than those provided by XENON100

