

HIGGS MASS AND STABILITY OF THE EW VACUUM

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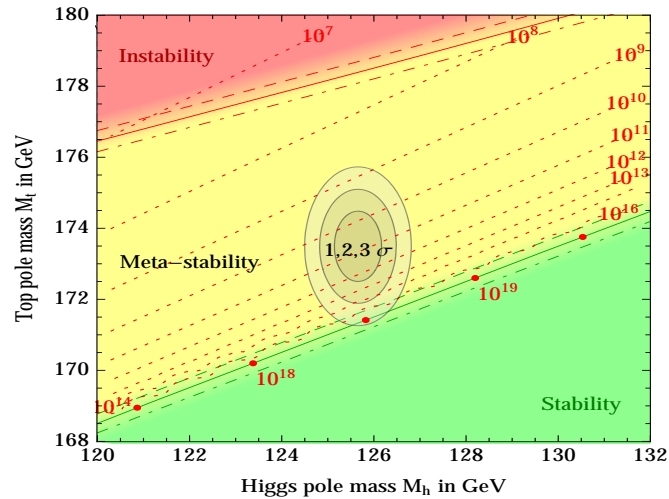
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Introduction and Summary

Stability / Metastability : very delicate issue...



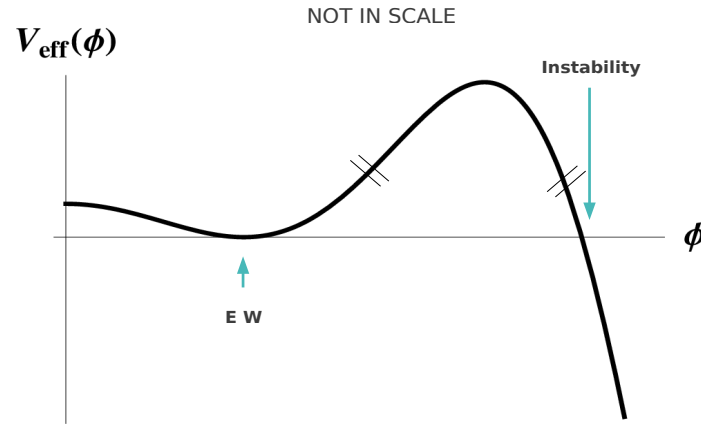
Degrassi et al., JHEP 2012 ; Buttazzo et al., arXiv:1307.3536 [hep-ph]

A complete analysis does not seem to lead to this phase diagram

(Why?)

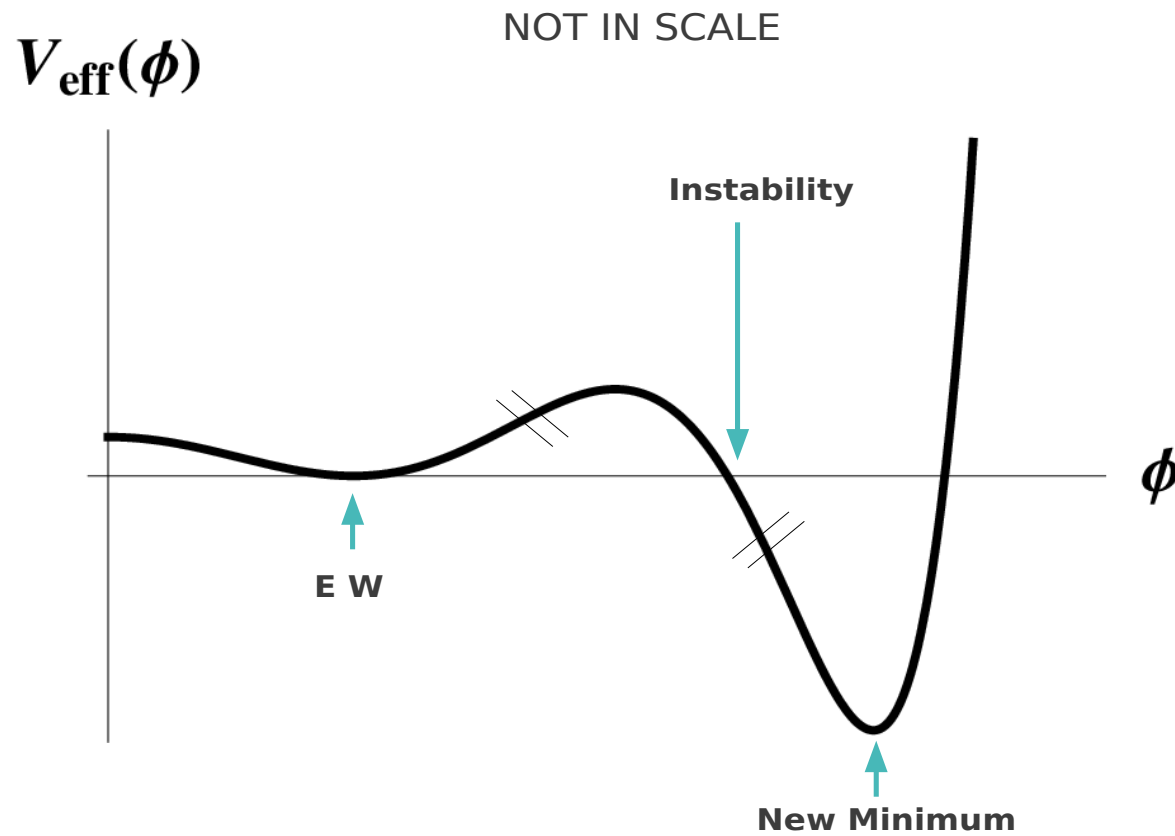
... We learn a quite different lesson ... Interesting consequences ?

One-Loop Effective Potential $V^{1l}(\phi)$



$$\begin{aligned}
 V^{1l}(\phi) = & \frac{1}{2}m^2\phi^2 + \frac{\lambda}{24}\phi^4 + \frac{1}{64\pi^2} \left[\left(m^2 + \frac{\lambda}{2}\phi^2\right)^2 \left(\ln\left(\frac{m^2 + \frac{\lambda}{2}\phi^2}{\mu^2}\right) - \frac{3}{2}\right) \right. \\
 & + 3 \left(m^2 + \frac{\lambda}{6}\phi^2\right)^2 \left(\ln\left(\frac{m^2 + \frac{\lambda}{6}\phi^2}{\mu^2}\right) - \frac{3}{2}\right) + 6 \frac{g_1^4}{16}\phi^4 \left(\ln\left(\frac{\frac{1}{4}g_1^2\phi^2}{\mu^2}\right) - \frac{5}{6}\right) \\
 & \left. + 3 \frac{(g_1^2 + g_2^2)^2}{16}\phi^4 \left(\ln\left(\frac{\frac{1}{4}(g_1^2 + g_2^2)\phi^2}{\mu^2}\right) - \frac{5}{6}\right) - 12 h_t^4\phi^4 \left(\ln\frac{g^2\phi^2}{\mu^2} - \frac{3}{2}\right) \right]
 \end{aligned}$$

RG Improved Effective Potential $V_{eff}(\phi)$



The instability occurs for **large values** of the field

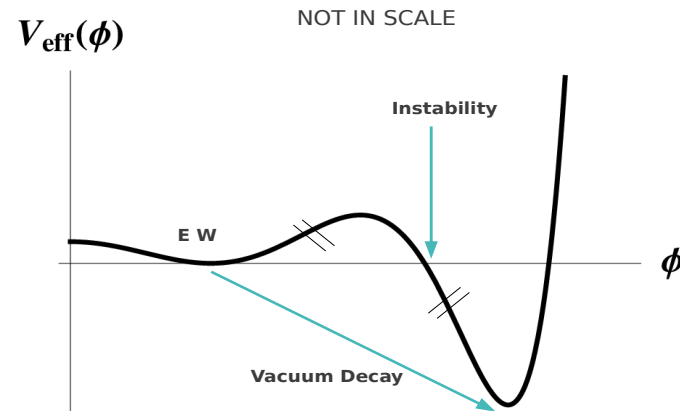
⇒ $V_{eff}(\phi)$ well approximated by keeping only the quartic term :

$$V_{eff}(\phi) \sim \frac{\lambda_{eff}(\phi)}{24} \phi^4$$

$\lambda_{eff}(\phi)$ depends on ϕ essentially as $\lambda(\mu)$ depends on μ

⇒ Read the **Effective Potential** from the $\lambda(\mu)$ flow

Metastability Scenario



Tunnelling between the **Metastable EW Vacuum** and the **True Vacuum**.
As long as **EW vacuum lifetime** larger than the **age of the Universe**

.... we may well live in the **Meta-Stable (EW) Vacuum**

How is the EW vacuum lifetime (**tunneling time**) computed?

Tunnelling time τ

$$\frac{1}{\tau} = T_U^3 \frac{S[\phi_b]^2}{4\pi^2} \left| \frac{\det' [-\partial^2 + V''(\phi_b)]}{\det [-\partial^2 + V''(v)]} \right|^{-1/2} e^{-S[\phi_b]}$$

$\phi_b(r)$: **Bounce Solution to the Euclidean Equation of Motion**

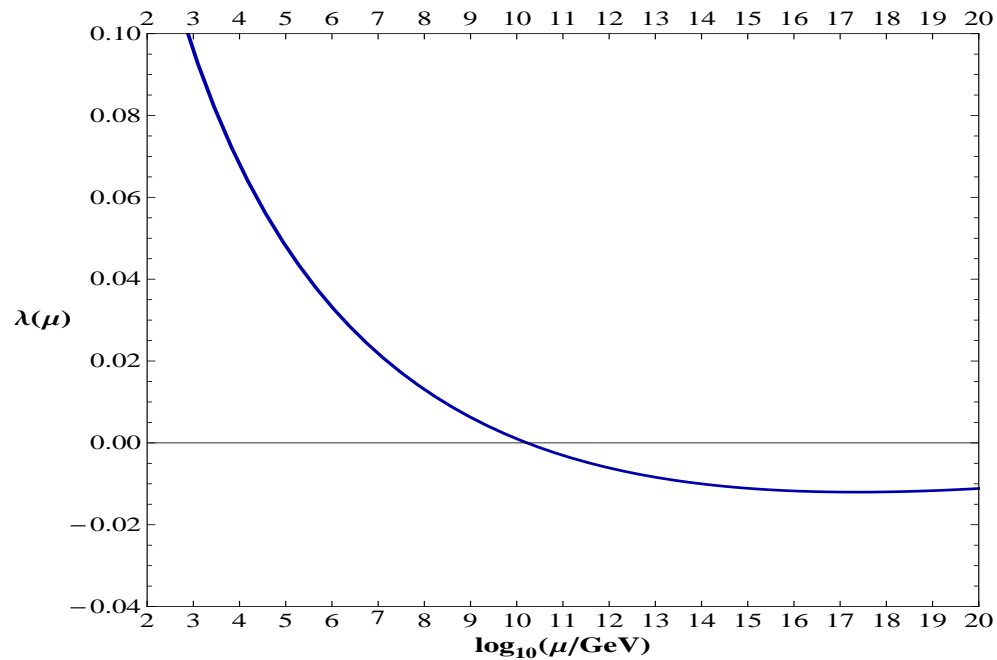
$$r^2 = x_\mu x_\mu$$

S. Coleman, Phys. Rev. D 15 (1977) 2929

C.G.Callan, S.Coleman, Phys. Rev. D 16 (1977) 1762

Tunnelling and Bounces

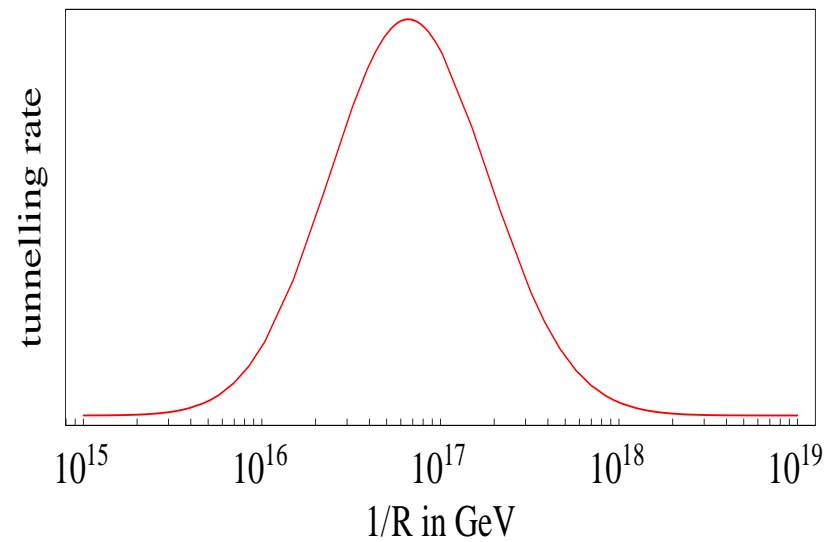
Bounce solutions to the euclidean equation of motion for $V(\phi) = \frac{\lambda}{4}\phi^4$ with constant negative λ , a good approximation in this range!



Bounces : $\phi_b(r) = \sqrt{\frac{2}{|\lambda|}} \frac{2R}{r^2+R^2}$, $S[\phi_b] = \frac{8\pi^2}{3|\lambda|}$

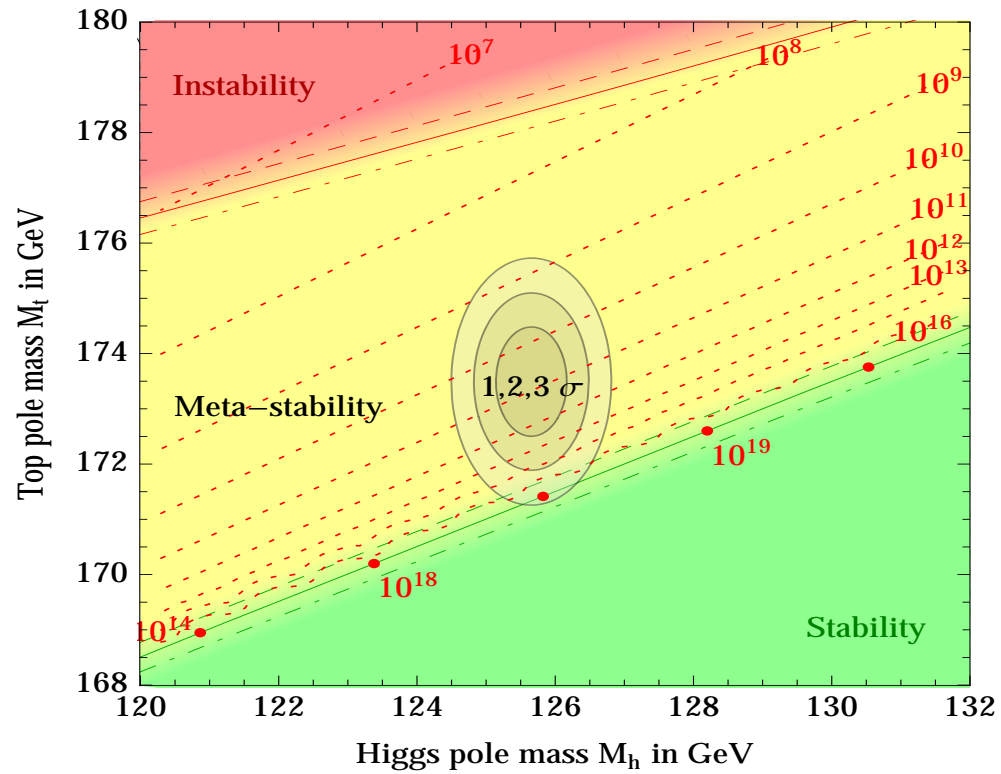
R : size of the bounce - **Degeneracy at the Classical Level**

Degeneracy removed at the Quantum Level



G. Isidori, G. Ridolfi, A. Strumia, Nucl.Phys.B 609 (2001) 387

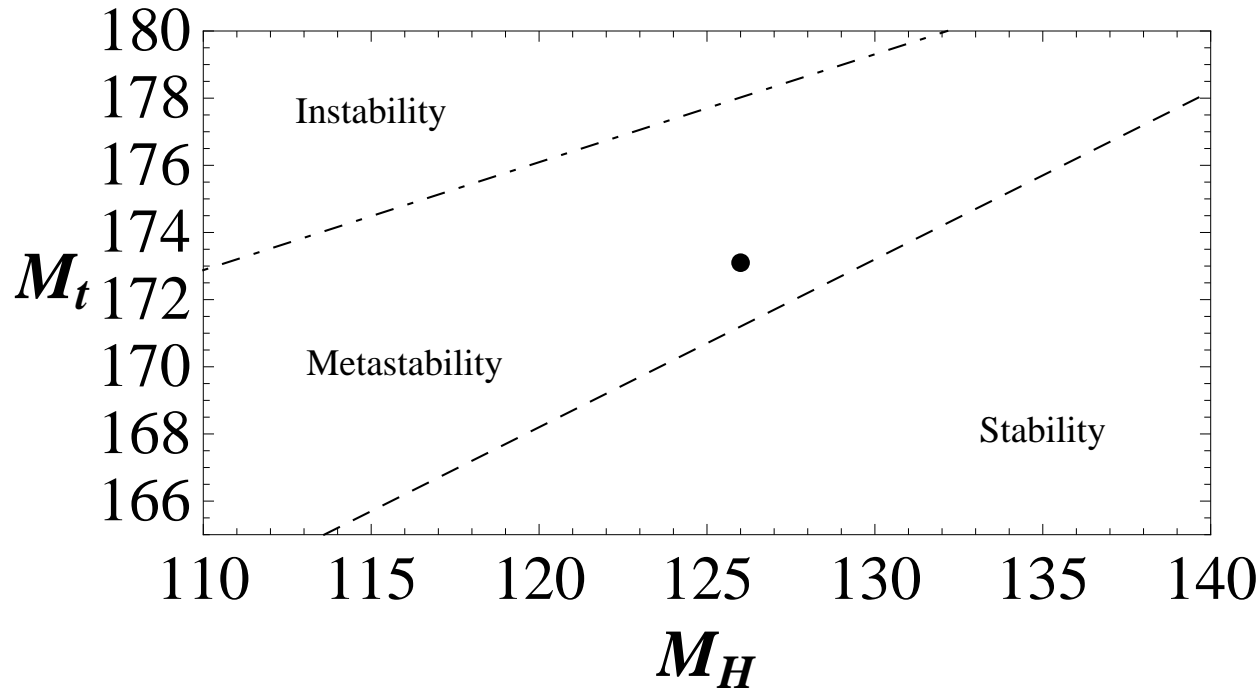
Phase Diagram in the $M_H - M_t$ plane



Buttazzo, Degrandi, Giardino, Giudice, Sala, Salvio, Strumia arXiv:1307.3536 [hep-ph]

Phase Diagram in the $M_H - M_t$ plane

V.B. , E. Messina, arXiv:1307.5193 [hep-ph]

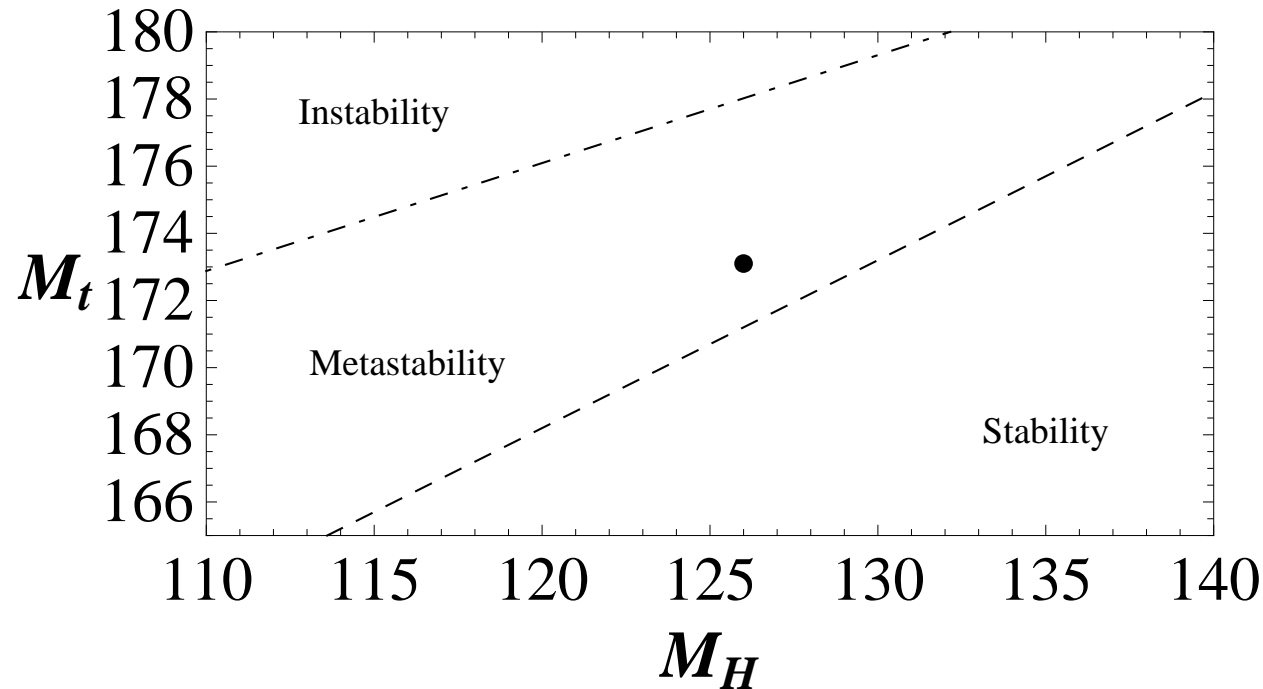


Stability region : $V_{eff}(v) < V_{eff}(\phi_{min}^{(2)})$. *Meta-stability* region : $\tau > T_U$.

Instability region : $\tau < T_U$. Dashed line : $V_{eff}(v) = V_{eff}(\phi_{min}^{(2)})$.

Dashed - dotted line : M_H and M_t such that $\tau = T_U$.

For $M_t \sim 173.1 \text{ GeV}$, $M_H \sim 126 \text{ GeV}$: SM within Metastability Region



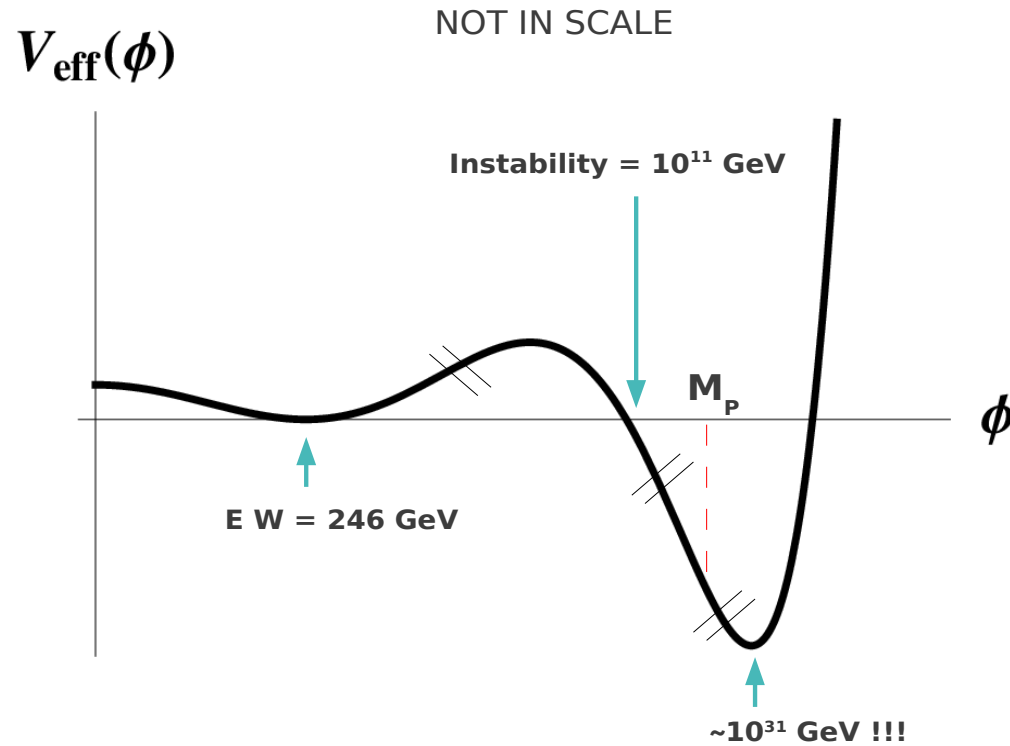
Conclusion(s) : This phase-diagram shows that for present values of M_H and M_t :

- Standard Model valid all the way up to the Planck scale ...
- We are close to criticality ...

..... **However**

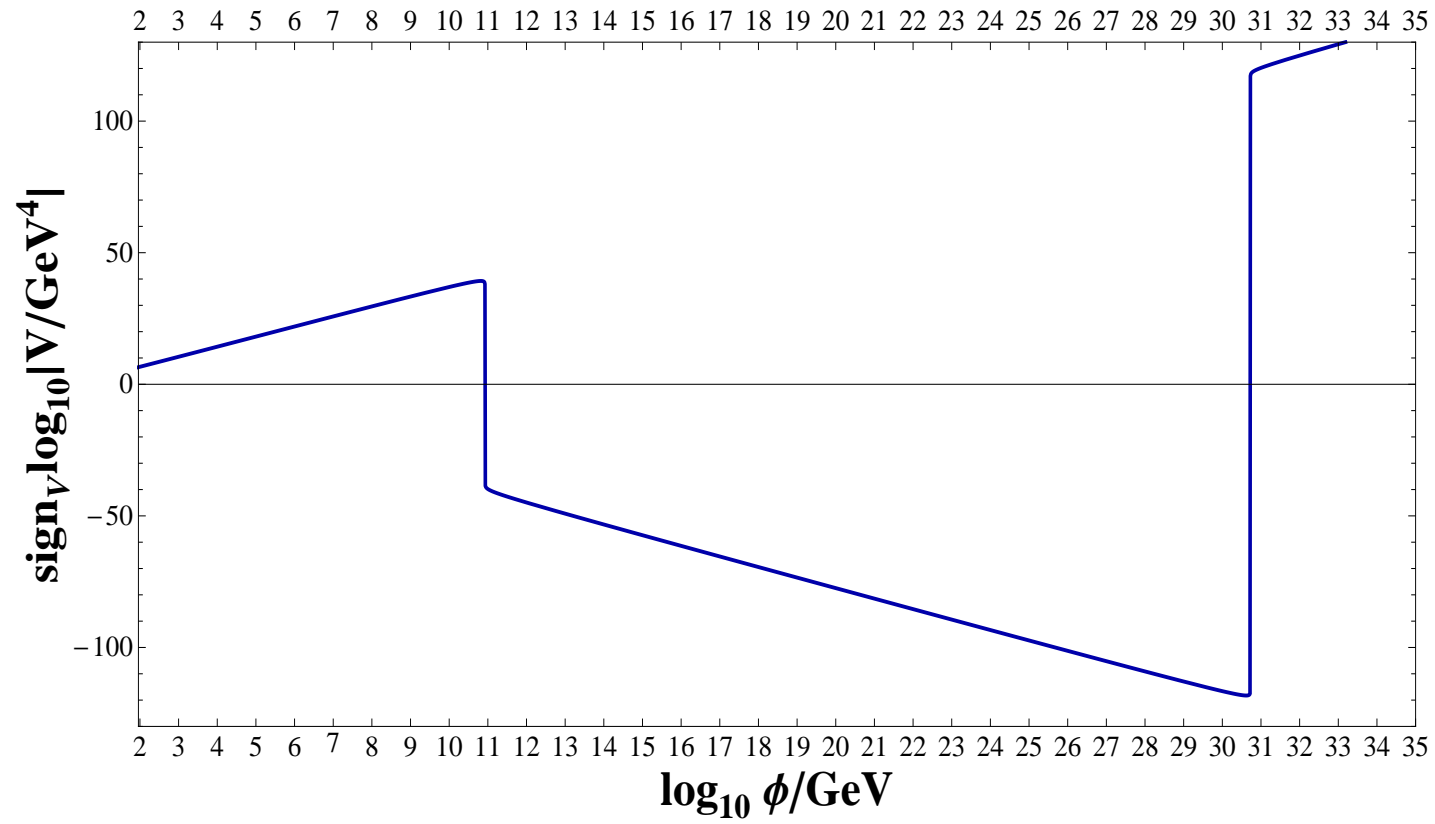
$$M_H = 126 \text{ GeV}$$

$$M_t = 173.1 \text{ GeV}$$



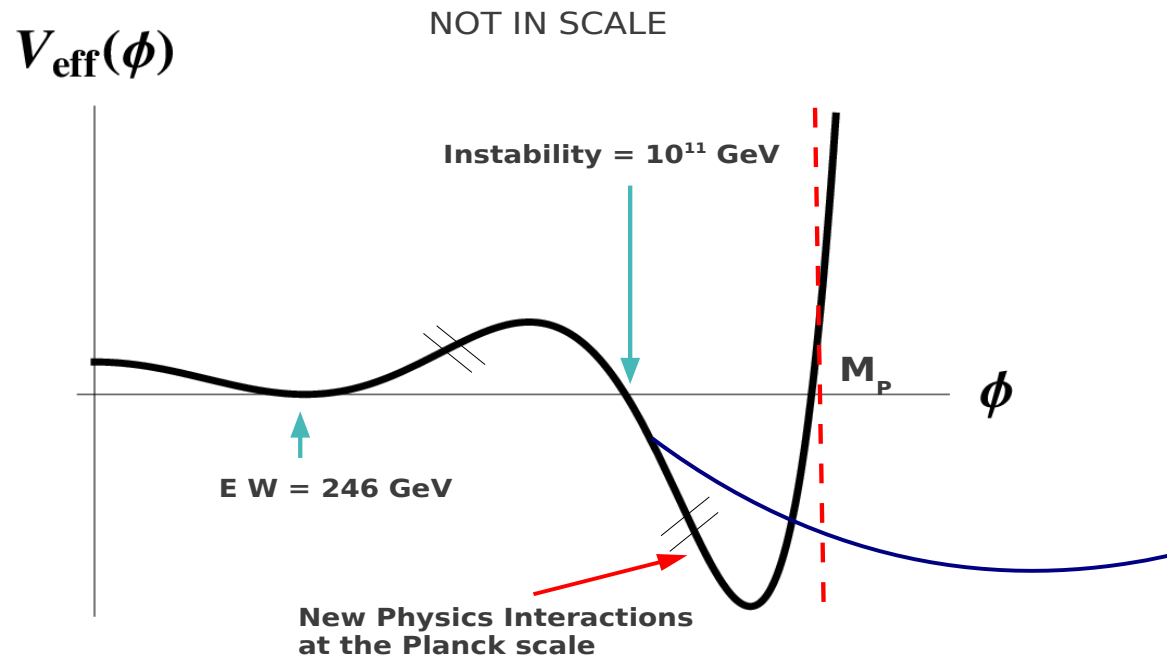
New minimum at $\phi \sim 10^{31}$ GeV !!!! Now picture in scale

Effective Potential ($M_H = 126 \text{ GeV}$, $M_t = 173.1 \text{ GeV}$)



Log-Log

New Physics Interactions stabilize the potential around the Planck scale



Are New Physics Operators harmless for computing τ ?
... It was thought they are ... But ...

... New Physics Interactions at the Planck scale ...

Add ϕ^6 and ϕ^8 to the SM Higgs potential:

$$V(\phi) = \frac{\lambda}{4}\phi^4 + \frac{\lambda_6}{6} \frac{\phi^6}{M_P^2} + \frac{\lambda_8}{8} \frac{\phi^8}{M_P^4}$$

The Effective Potential $V_{eff}^{new}(\phi)$ is modified as :

$$V_{eff}^{new}(\phi) = V_{eff}(\phi) + \frac{\lambda_6(\phi)}{6M_P^2} \xi(\phi)^6 \phi^6 + \frac{\lambda_8(\phi)}{8M_P^4} \xi(\phi)^8 \phi^8$$

For $\phi < M_P$: $V_{eff}^{new}(\phi)$ **practically coincides with** $V_{eff}(\phi)$

For $\phi \sim M_P$: $V_{eff}^{new}(\phi)$ **depends on** λ_6 and λ_8 ...

Two different representative cases

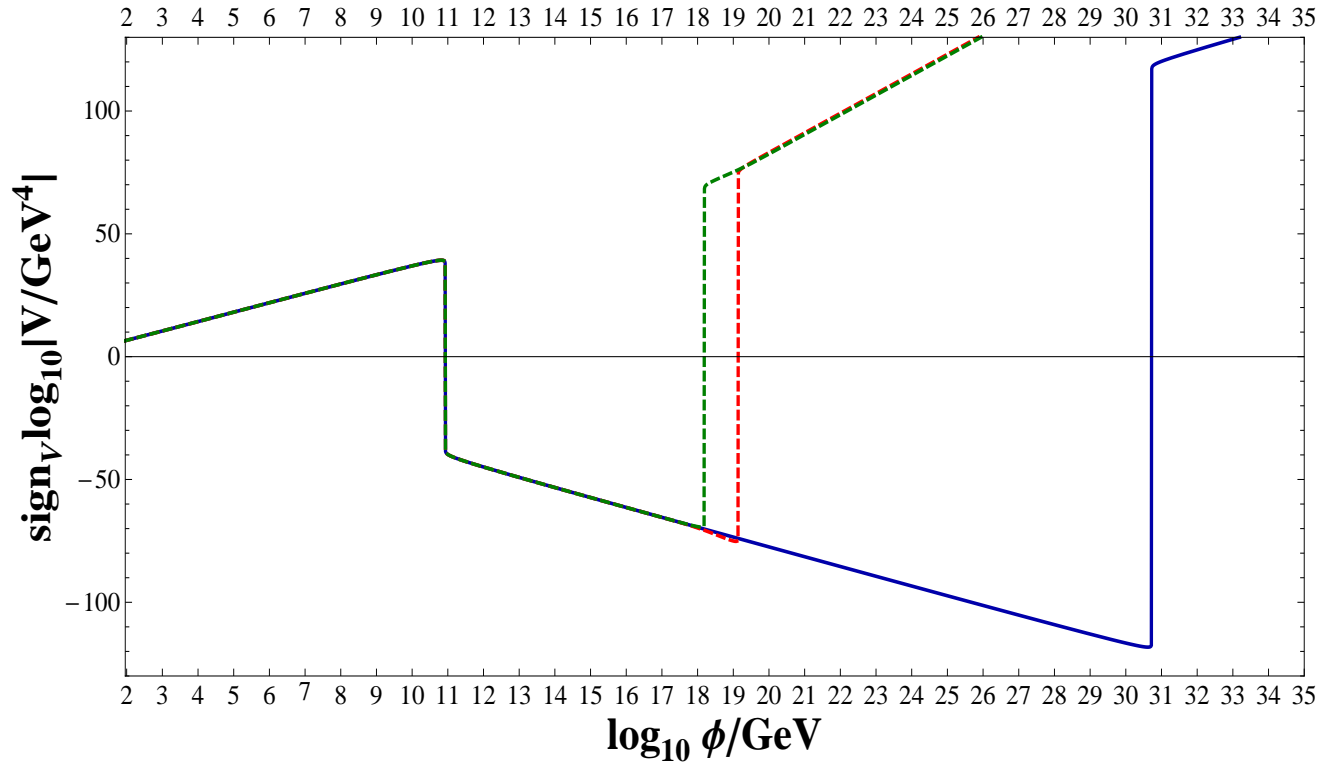
1. λ_6 negative and λ_8 positive

Example : $\lambda_6(M_P) = -2$ $\lambda_8(M_P) = 2.1$

2. both λ_6 and λ_8 positive

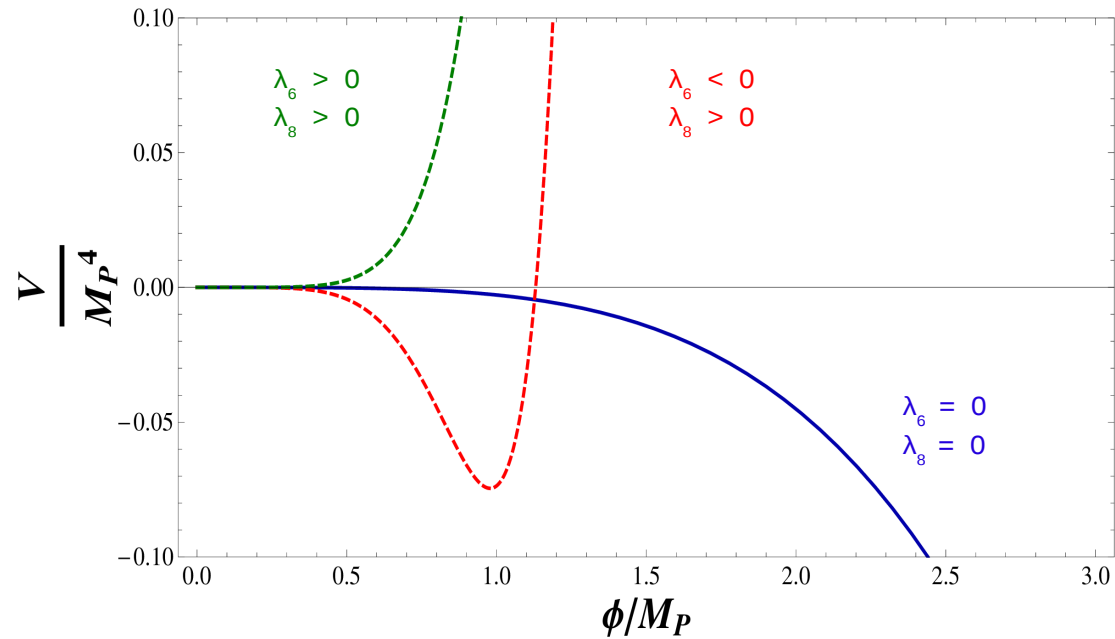
Example : $\lambda_6(M_P) = 1$ $\lambda_8(M_P) = 0.5$

Effective Potential $M_H = 126$ $M_t = 173.1$ **Log-Log Plot**



- Blue line :** $V_{eff}(\phi)$ no higher order terms
- Red line :** $V_{eff}^{new}(\phi)$ with $\lambda_6(M_P) = -2$ $\lambda_8(M_P) = 2.1$
- Green line :** $V_{eff}^{new}(\phi)$ with $\lambda_6(M_P) = 1$ $\lambda_8(M_P) = 0.5$

Zoom around the Planck scale



- | | | |
|---------------------|-----------------------|---------------------------------------------------|
| Blue line : | $V_{eff}(\phi)$ | no higher order terms |
| Red line : | $V_{eff}^{new}(\phi)$ | with $\lambda_6(M_P) = -2$ $\lambda_8(M_P) = 2.1$ |
| Green line : | $V_{eff}^{new}(\phi)$ | with $\lambda_6(M_P) = 1$ $\lambda_8(M_P) = 0.5$ |

1. The case $\lambda_6 < 0$ ($\lambda_6 = -2$) $\lambda_8 > 0$ ($\lambda_8 = 2.1$)

a. Up to $\eta \simeq 0.78M_P$, V_{eff}^{new} very well approximated by :

$$V_{eff}^{new}(\phi) = \frac{\lambda_{eff}}{4} \phi^4$$

$$\lambda_{eff} = \lambda + \frac{2}{3} \lambda_6 \frac{\eta^2}{M_P^2} + \frac{1}{2} \lambda_8 \frac{\eta^4}{M_P^4} \simeq -0.437.$$

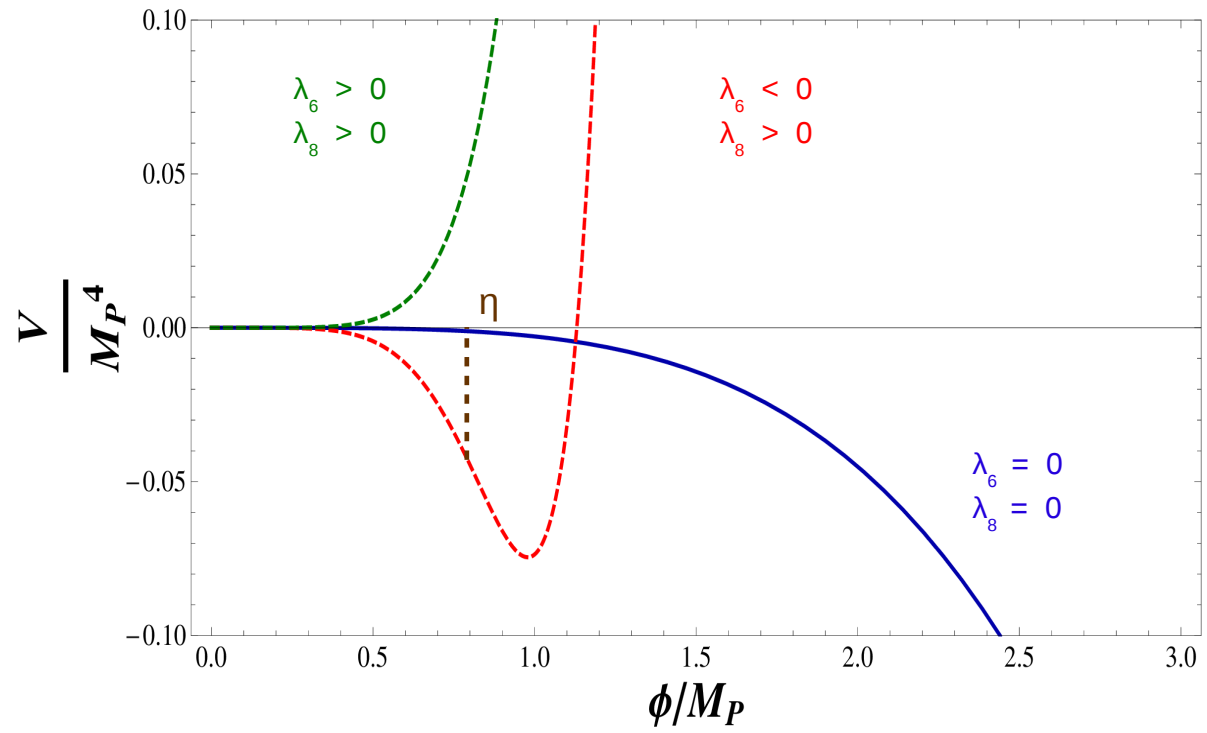
b. For $\phi \gtrsim \eta$, $V_{eff}^{new}(\phi)$ can be linearized :

$$V(\phi) = \frac{\lambda_{eff}}{4} \eta^4 - \frac{\lambda_{eff} \eta^3}{\gamma} (\phi - \eta)$$

$$\gamma = -\lambda_{eff} \eta^3 \left(\lambda \eta^3 + \lambda_6 \frac{\eta^5}{M_P^2} + \lambda_8 \frac{\eta^7}{M_P^4} \right)^{-1}$$

This is all what we need to know to compute τ !!!

K. Lee, E.J. Weinberg, Nucl. Phys. B 267 (1986) 181



Equation of motion admits **Two Types (!)** of bounce solutions:

Type 1 (Old Friend !!!) : $\phi_b^{(1)}(r) = \sqrt{\frac{2}{|\lambda_{eff}|}} \frac{2R}{r^2 + R^2}$

R : size of these bounces

Action degenerate with R : $S[\phi_b^{(1)}] = \frac{8\pi^2}{3|\lambda_{eff}|}$

Type 2: (New Comer !!!) $\phi_b^{(2)}(r) = \begin{cases} 2\eta - \eta^2 \sqrt{\frac{|\lambda_{eff}|}{8}} \frac{r^2 + \bar{R}^2}{\bar{R}} & 0 < r < \bar{r} \\ \sqrt{\frac{8}{|\lambda_{eff}|}} \frac{\bar{R}}{r^2 + \bar{R}^2} & r > \bar{r} \end{cases}$

$$\bar{r}^2 = \frac{8\gamma}{\lambda_{eff}\eta^2}(1 + \gamma) \quad , \quad \bar{R}^2 = \frac{8}{|\lambda_{eff}|} \frac{\gamma^2}{\eta^2} .$$

Only for

$$-1 < \gamma < 0$$

\bar{R} : size of the bounce

Action : $S[\phi_b^{(2)}] = (1 - (\gamma + 1)^4) \frac{8\pi^2}{3|\lambda_{eff}|}$

If New Physics Interactions not explicitly included

Only Type 1 Solution !!!

... Tunnelling time computed accordingly (literature !!!) ...

$$\frac{1}{\tau} = T_U^3 \frac{S[\phi_b]^2}{4\pi^2} \left| \frac{\det' [-\partial^2 + V''(\phi_b)]}{\det [-\partial^2 + V''(v)]} \right|^{-1/2} e^{-S[\phi_b]}$$

Even considering tree-level only (neglecting determinants)

As $S[\phi_b^{(1)}] = \frac{8\pi^2}{3|\lambda_{eff}|}$ and $S[\phi_b^{(2)}] = (1 - (\gamma + 1)^4) \frac{8\pi^2}{3|\lambda_{eff}|}$

When Type 2 solution exists ($-1 < \gamma < 0$) : contribution from Type 1 bounces exponentially suppressed (!!!) with respect to contribution from Type 2 bounce

... Some numbers ...

Our case : $\lambda_6(M_P) = -2$ $\lambda_8(M_P) = 2.1$ ($\gamma \simeq -0.963$)

$$\frac{1}{\tau} = T_U^3 \frac{S[\phi_b^{(2)}]^2}{4\pi^2} \left| \frac{\det' [-\partial^2 + V''(\phi_b)]}{\det [-\partial^2 + V''(v)]} \right|^{-1/2} e^{-S[\phi_b^{(2)}]}$$

$$\tau \sim 10^{-219} T_U \lll T_U \quad \text{!!!!}$$

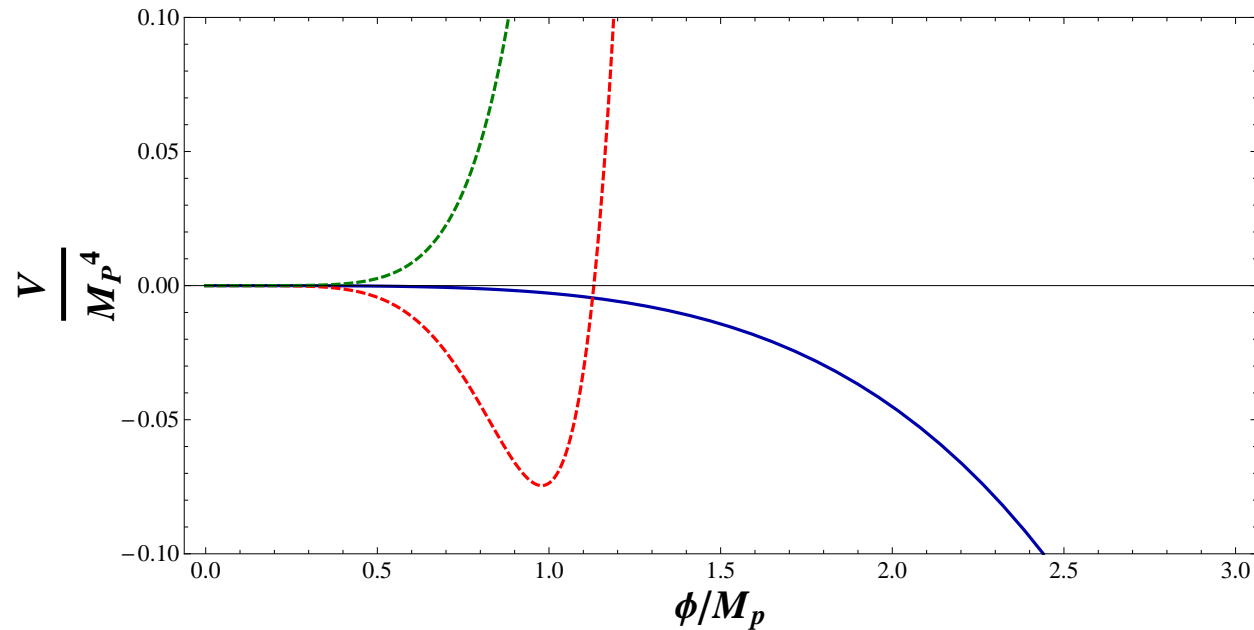
Literature case : $\lambda_6(M_P) = 0$ $\lambda_8(M_P) = 0$

$$\frac{1}{\tau} = T_U^3 \frac{S[\phi_b^{(1)}]^2}{4\pi^2} \left| \frac{\det' [-\partial^2 + V''(\phi_b)]}{\det [-\partial^2 + V''(v)]} \right|^{-1/2} e^{-S[\phi_b^{(1)}]}$$

$$\tau \sim 10^{544} T_U \ggg T_U \quad \text{!!!!}$$

New Physics Interactions at the Planck scale do matter !!!

2. Case $\lambda_6 > 0$ and $\lambda_8 > 0$ ($\gamma < -1$) ... only Type 1 bounces !!!



$$S[\phi_b^{(1)}] = \frac{8\pi^2}{3|\lambda_{eff}|} \quad S[\phi_b^{(2)}] = (1 - (\gamma + 1)^4) \frac{8\pi^2}{3|\lambda_{eff}|}$$

γ is the crucial parameter ! Two solutions only for :

$$-1 < \gamma < 0$$

Otherwise only Type 1 : $\tau = \tau_{SM}$!!!

Determinant contributions do not change these results !!!

Logarithm of the fluctuation determinant (Higgs sector only) :

$$\log \left(\frac{\det'(-\partial^2 + V''(\phi_b))}{\det(-\partial^2)} \right)^{1/2} = \frac{1}{2} \sum_{l=0}^{\infty} (l+1)^2 \ln \rho_l$$

$$\rho_l = \lim_{r \rightarrow \infty} \rho_l(r)$$

Each $\rho_l(r)$ solution of the differential equation:

$$\rho_l''(r) + \frac{(2l+d-1)}{r} \rho_l'(r) - V''(\phi_b(r)) \rho_l(r) = 0 \quad (\rho_l(0) = 1 ; \rho_l'(0) = 0)$$

\overline{MS} renormalized sum (excluding negative and zero modes):

$$\left[\frac{1}{2} \sum_{l>1}^{\infty} (l+1)^2 \ln \rho_l \right]_{Ren} = \frac{1}{2} \sum_{l>1}^{\infty} (l+1)^2 \ln \rho_l - \frac{1}{2} \sum_{l=0}^{\infty} (l+1)^2 \left[\frac{\int_0^{\infty} dr r V''}{2(l+1)} - \frac{\int_0^{\infty} dr r^3 (V'')^2}{8(l+1)^3} \right]$$

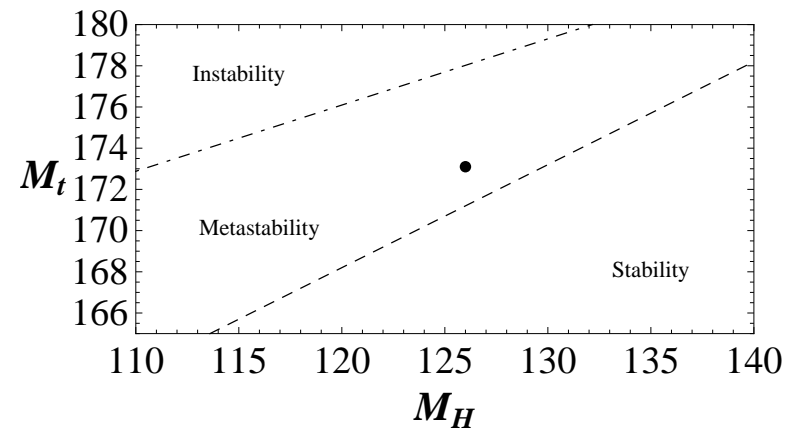
$$- \frac{1}{8} \int_0^{\infty} dr r^3 (V'')^2 \left[\ln \left(\frac{\mu r}{2} \right) + \gamma_E + 1 \right]$$

For our example ($\lambda_6(M_P) = -2$ $\lambda_8(M_P) = 2.1$), adding determinant contribution (Higgs sector only), including negative and zero modes :

$$\tau \sim 10^{-212} T_U \lll T_U \quad \text{!!!!}$$

Top , Gauge , Goldstone contributions do not change this result!

Phase Diagram in the $M_H - M_t$ plane (Literature)



For $M_H \sim 126$ GeV and $M_t \sim 173.1$ GeV

Neglecting New Interactions at the Planck scale

EW vacuum well inside Metastability Region, close to Stability line

Extremely long-lived Metastable State !

This is why often stated :

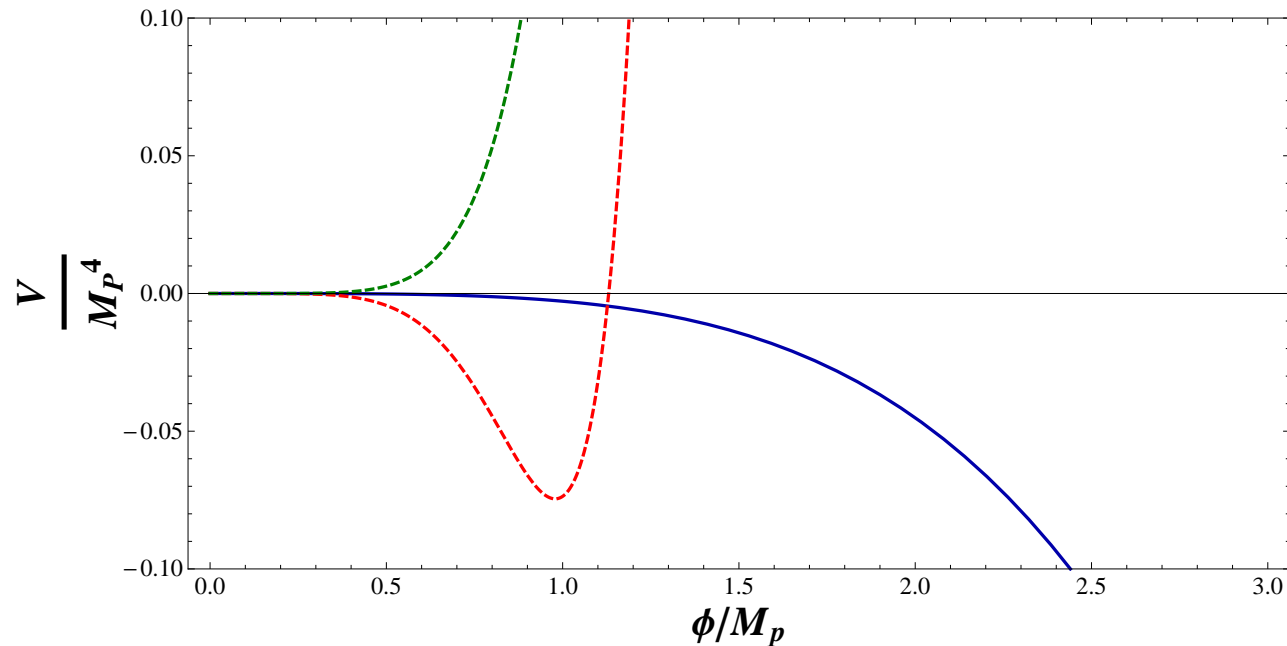
The SM is an effective theory that may be valid all the way up to M_P !

SM : effective theory valid up to M_P ???

Lesson : This is not generically true !!!

New physics interactions at the Planck scale may turn the EW vacuum from a **very long-lived metastable state** to a **highly unstable state**.

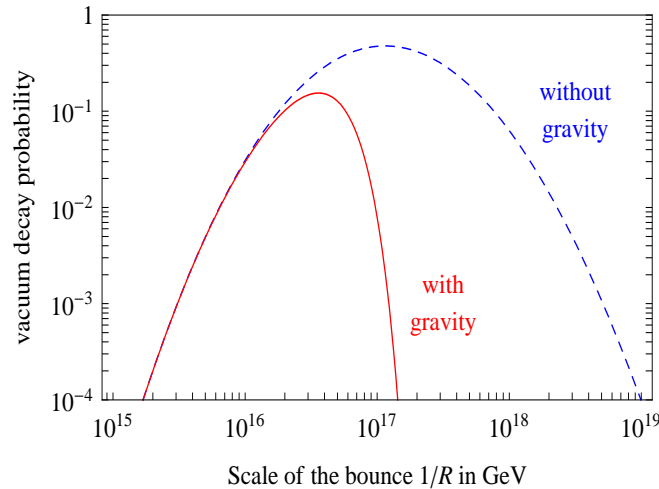
When $V_{eff}^{new}(\phi)$ lies above $V_{eff}(\phi)$, τ is not affected by new physics. On the contrary, when $V_{eff}^{new}(\phi)$ lies below $V_{eff}(\phi)$, New Physics Interaction at the Planck scale have strong impact on τ , turning $\tau \gg T_U$ to $\tau \ll T_U$!



Conclusions

V.B. , E. Messina, arXiv:1307.5193 [hep-ph]

- **Lifetime τ of the EW vacuum** strongly depends on New Physics Interactions at the Planck scale
- **SM Phase Diagram** strongly depends on New Physics Interactions at the Planck scale
- **Metastability Scenario** (based on the assumption that τ does not depend on New Physics Interactions at the Planck scale) has to be reconsidered
- These results provide **constraints on New Physics beyond SM**
- A similar analysis can be done also if the new physics scale lies **below the Planck scale**
- Analysis also relevant for : **Higgs potential with two degenerate minima, $\lambda(M_P) \sim 0$, $\beta(\lambda(M_P)) \sim 0$, Higgs driven inflation...** In all of these cases, the relevant physical scale **dangerously close to the Planck scale** \rightarrow high sensitivity to New Physics Interactions



Isidori, Rychkov, Strumia, Tetradis, “Gravitational corrections to Standard Model vacuum decay” ,
 Phys.Rev.D77:025034,2008

... Adding to the SM action possible dimension-6 non-renormalizable operators suppressed by the Planck scale \Rightarrow corrections to the bounce action. In particular, adding to the SM Lagrangian the operators

$$\Delta_6 = \frac{1}{M_{\text{Pl}}^2} \left(-\xi M_{\text{Pl}}^2 R |H|^2 + c_1 \frac{|H|^6}{3!} + c_2 |H|^2 |D_\mu H|^2 \right)$$

where ξ and $c_{1,2}$ are unknown dimensionless coefficients, gives the following correction

$$\Delta S'_{\text{gravity}} = \frac{8\pi^2}{15(M_{\text{Pl}} R \lambda)^2} \left(128\pi\xi + \frac{c_1}{|\lambda|} + 4c_2 \right) \dots$$

